



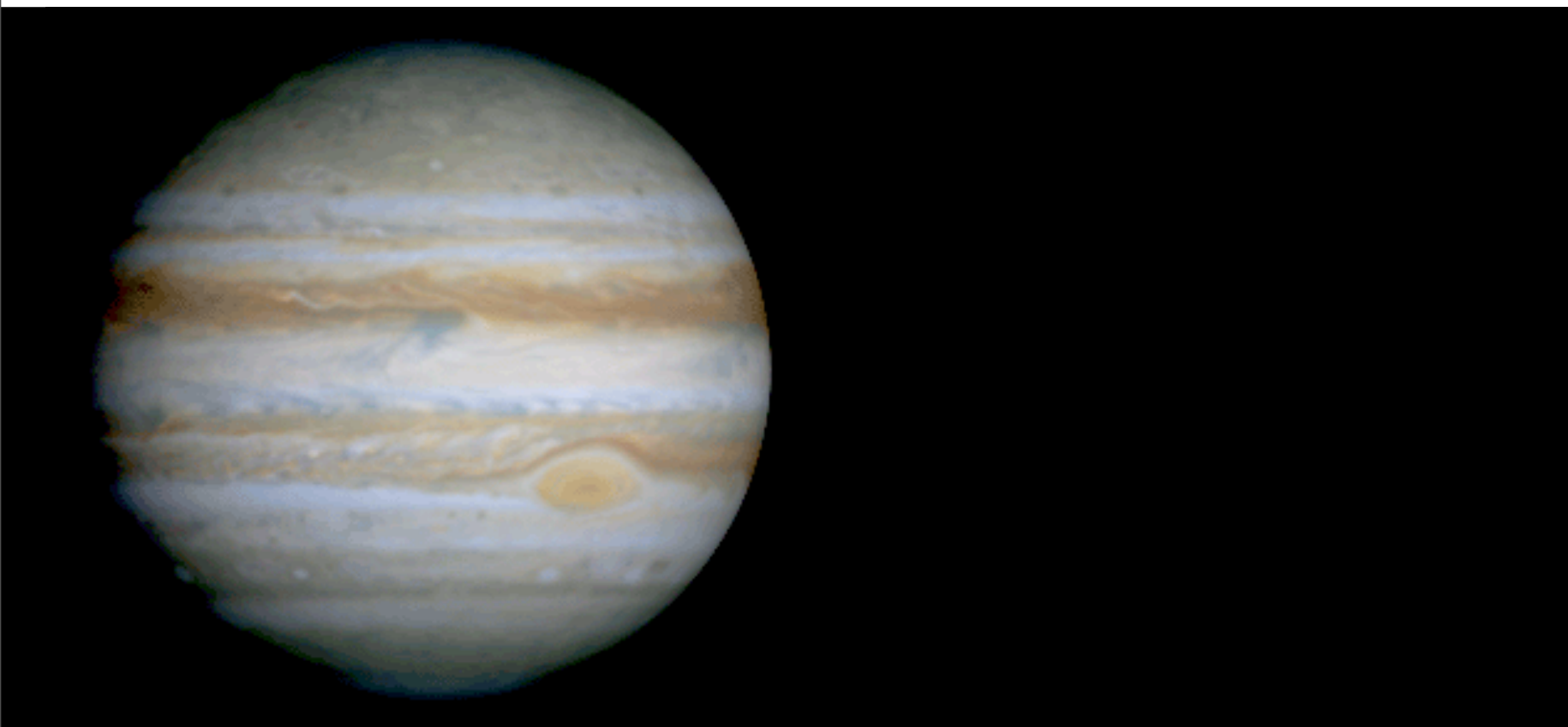
Verification of the predictions of SSST in nonlinear simulations

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National and Kapodistrian University of Athens

ISSI Group Meeting
April 2013



zonal flows coexist with turbulence

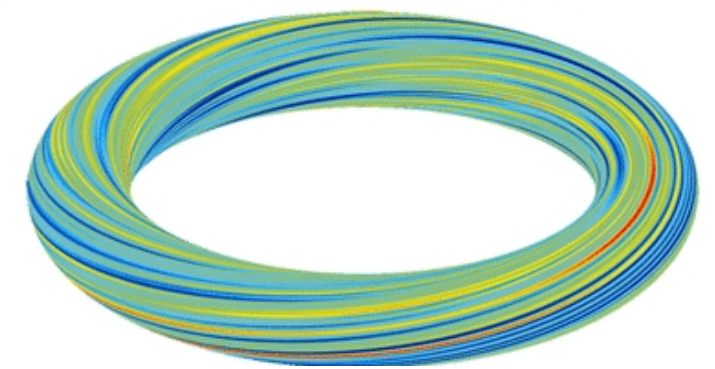
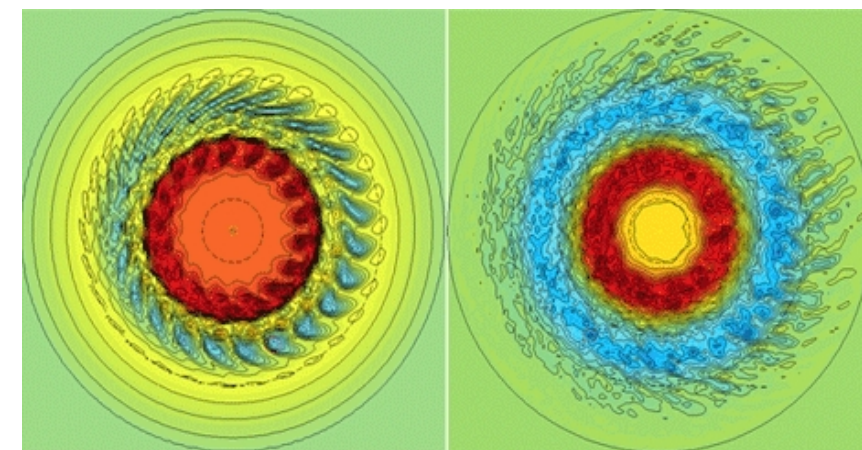


banded Jovian jets

NASA/Cassini Jupiter Images

polar front jet

NASA/Goddard Space Flight Center

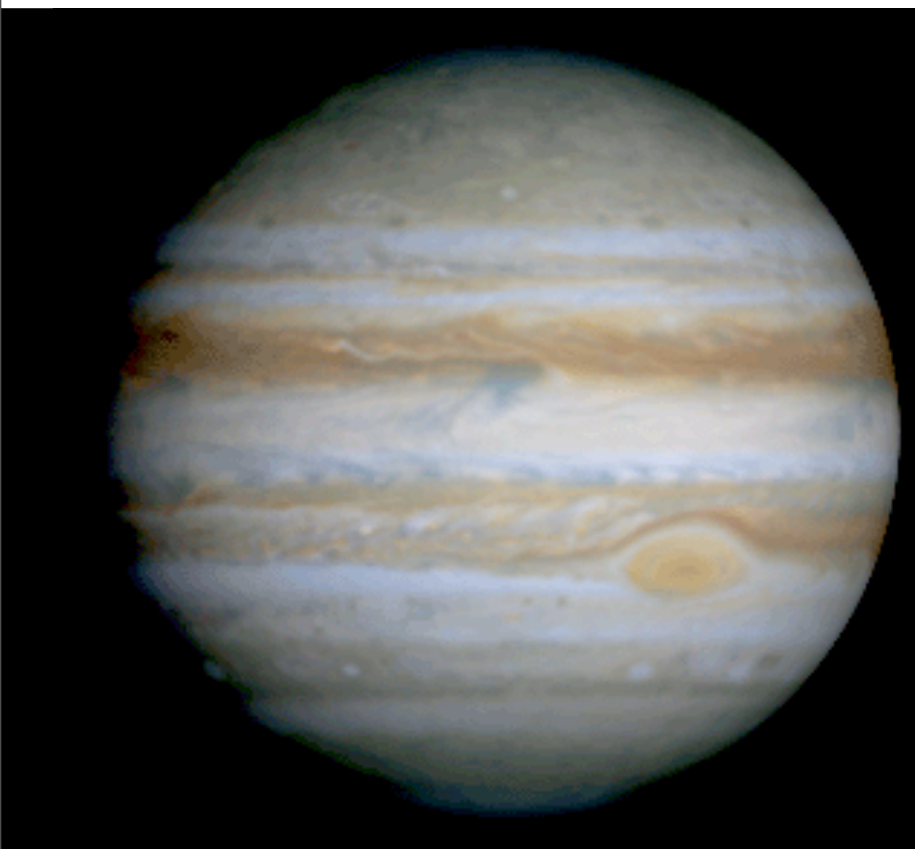


jets in tokamaks

courtesy: L.Villard

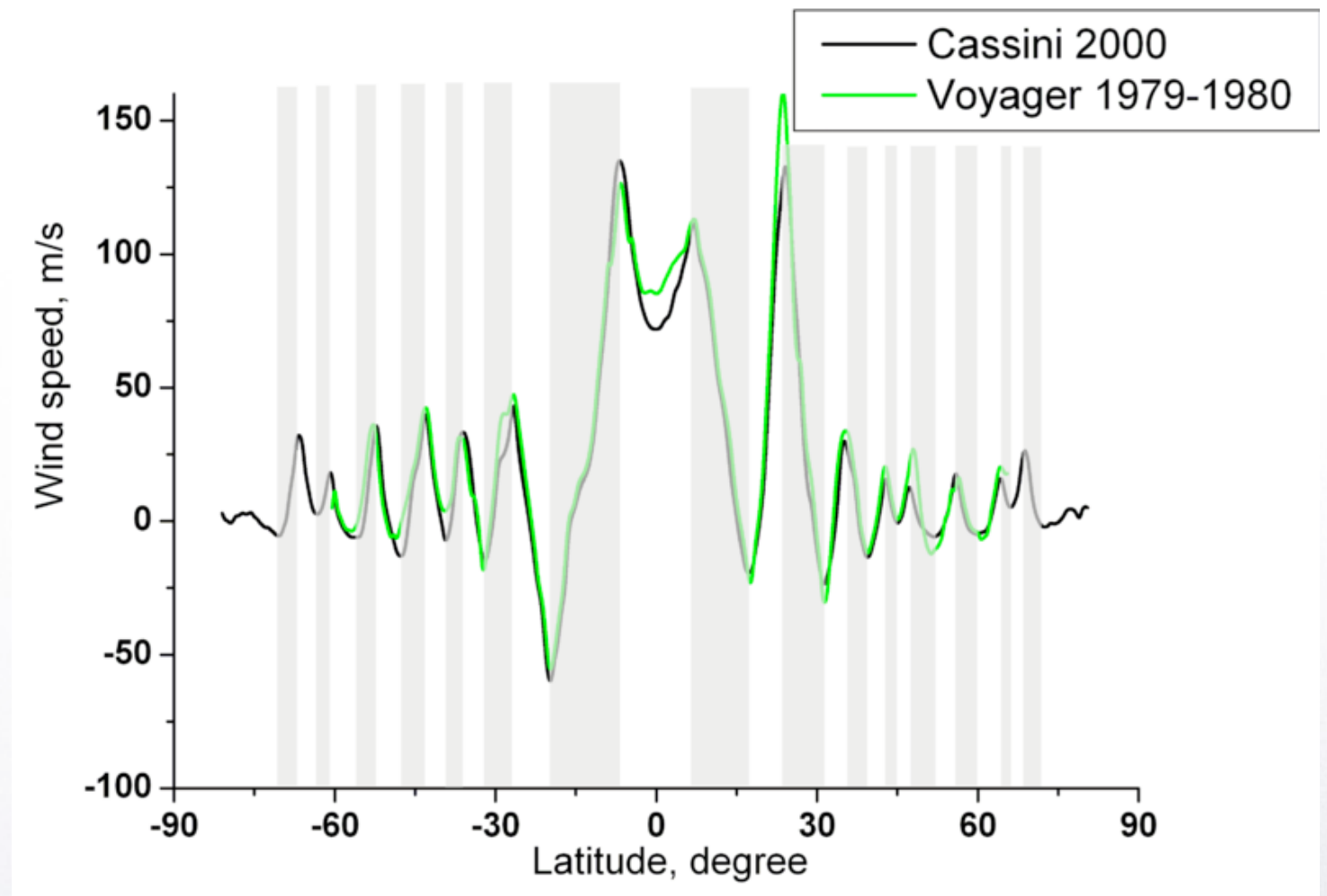


zonal flows coexist with turbulence



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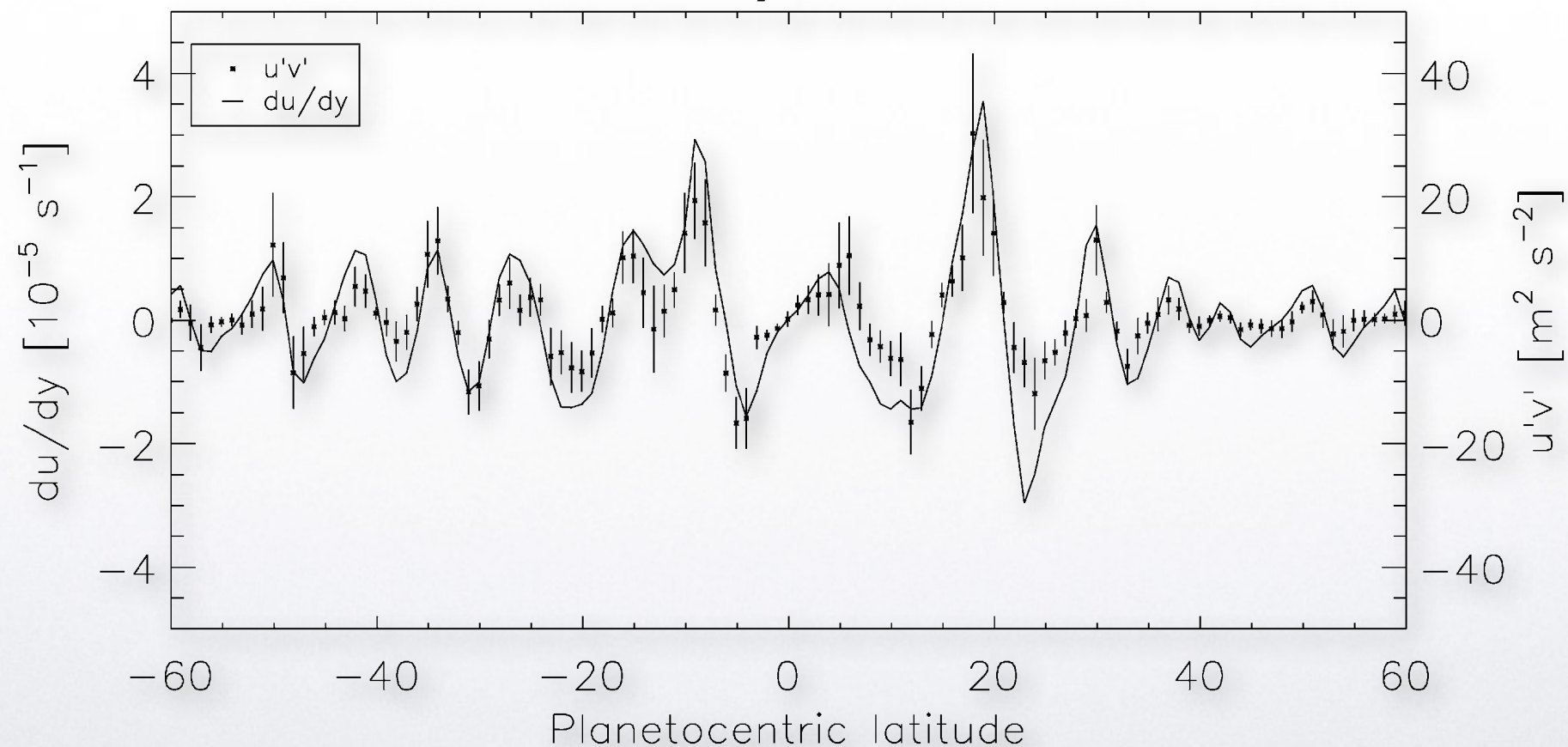
observed Jovian zonal winds
at cloud level

Vasadava & Showman, 2005



zonal flows are maintained by eddies

$$\frac{d}{dt} \int \frac{U^2}{2} dy = \int \frac{dU}{dy} \overline{u'v'} dy - \text{Dissipation}$$



(Salyk et. al. 2006)



Zonal - Eddy field decomposition

$$\varphi(x, y, t) = \Phi(y, t) + \varphi'(x, y, t)$$

where

$$\Phi(y, t) = \overline{\varphi}(y, t) = \frac{1}{L_x} \int_0^{L_x} \varphi(x', y, t) dx'$$



Zonal - Eddy field decomposition

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↑
zonal mean

↑
eddy

where

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SSST System

$$\frac{dU}{dt} = \sum_{k=1}^{N_k} \frac{k}{2} \text{Im} [\text{vecd} (\Delta_k^{-1} \mathbf{C}_k)] - r_m U$$

$$\frac{d\mathbf{C}_k}{dt} = \mathbf{A}_k(U) \mathbf{C}_k + \mathbf{C}_k \mathbf{A}_k(U)^\dagger + \epsilon \mathbf{Q}_k$$

admits equilibria $(U^E, \mathbf{C}_1^E, \dots, \mathbf{C}_{N_k}^E) \equiv (U^E, \mathbf{C}_k^E)$

homogeneous equilibrium: $(U^E = 0, \mathbf{C}_k^E = \epsilon \mathbf{Q}_k^E / 2r)$



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admits equilibria $(U^E, \mathbf{C}^E)$ N_y discretization points and N_k zonal harmonics $\Rightarrow$
 $\Rightarrow (U, \mathbf{C}_k)$ is a $(2N_k N_y^2 + N_y)$ state vector

homogeneous equilibrium

(e.g. $N_y = 128, N_k = 10 \Rightarrow (U, \mathbf{C}_k) \sim 3.3 \times 10^5$)



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Stability of SSST equilibrium

perturbing the SSST equilibrium $(U^E + \delta U, \mathbf{C}_k^E + \delta \mathbf{C}_k)$

$$\frac{d\delta U}{dt} = \sum_{k=1}^{N_k} \frac{k}{2} \text{Im} [\text{vecd} (\Delta_k^{-1} \delta \mathbf{C}_k)] - r_m \delta U$$

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$$\text{with } \delta \mathbf{A}_k \equiv \mathbf{A}_k(U^E + \delta U) - \mathbf{A}_k(U^E)$$

split for real/imaginary part and

search for eigensolutions: $(\delta \hat{U}, \delta \hat{\mathbf{C}}_{k,R}, \delta \hat{\mathbf{C}}_{k,I}) e^{\sigma t}$ for $k = 1, \dots, N_k$,

$$\sigma \begin{pmatrix} \delta \hat{U} \\ \delta \hat{\mathbf{C}}_{k,R} \\ \delta \hat{\mathbf{C}}_{k,I} \end{pmatrix} = \mathbb{L} \begin{pmatrix} \delta \hat{U} \\ \delta \hat{\mathbf{C}}_{k,R} \\ \delta \hat{\mathbf{C}}_{k,I} \end{pmatrix}$$



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Eigenanalysis of stability operator

for $(U^E = 0, C_k^E = \epsilon Q_k^E/2r) \Rightarrow \delta\hat{U} = e^{iny}$

1. assume $\sigma, \delta\hat{U}_n \Rightarrow \delta\hat{C}_{k,n} \Rightarrow \delta(\overline{v'\zeta'})_n$

2. use mean flow perturbation equation

$$\sigma\delta\hat{U}_n = \delta(\overline{v'\zeta'})_n - r_m\delta\hat{U}_n$$

to calculate new σ

3. repeat until convergence

(this method can be generalized for $U^E \neq 0$ equilibria)

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this method works extremely well and allows us to attack the $(2N_k N_y^2 + N_y) \times (2N_k N_y^2 + N_y)$ eigenvalue problem which would otherwise be impossible

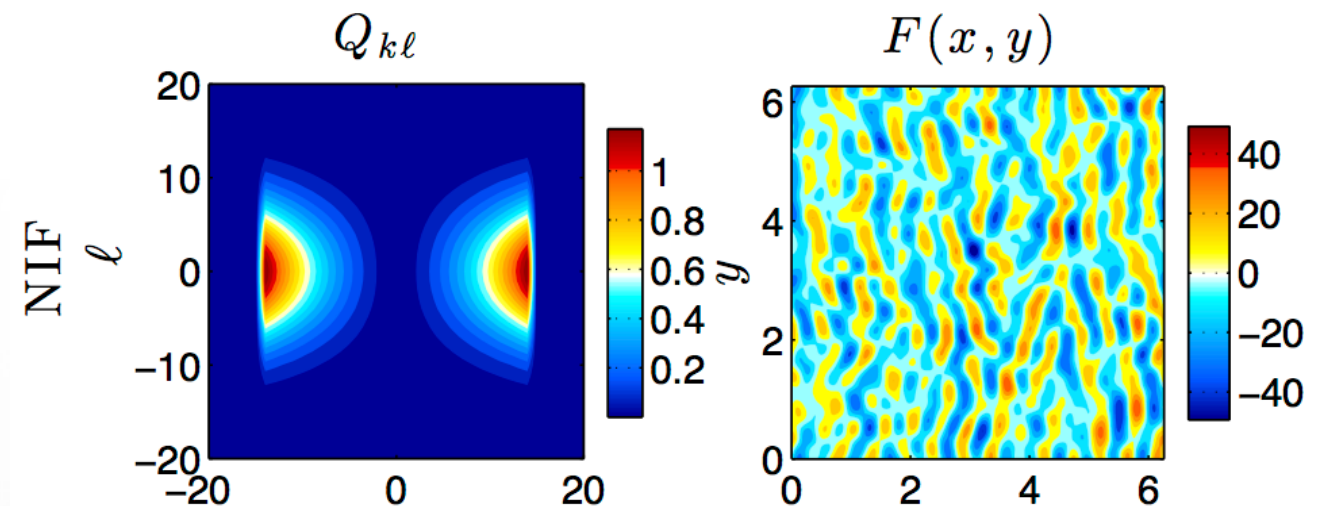
(this method can be generalized for $U^E \neq 0$ equilibria)



Types of stochastic forcing: NIF & IRFh

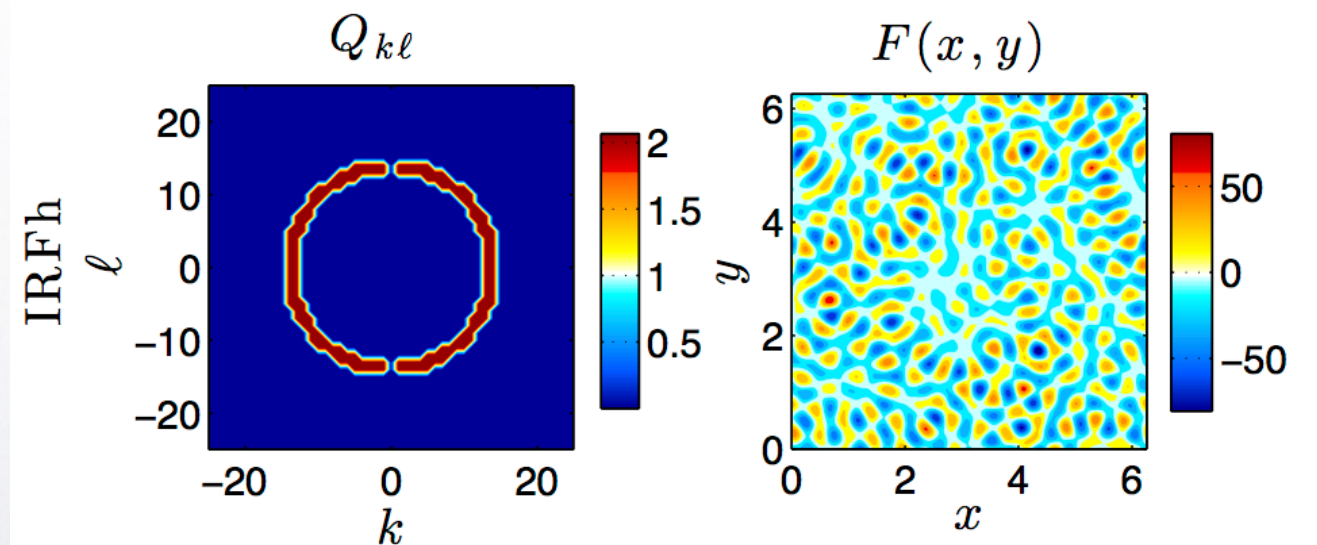
Non-isotropic Forcing (NIF)

parametrize excitation by
baroclinic instabilities



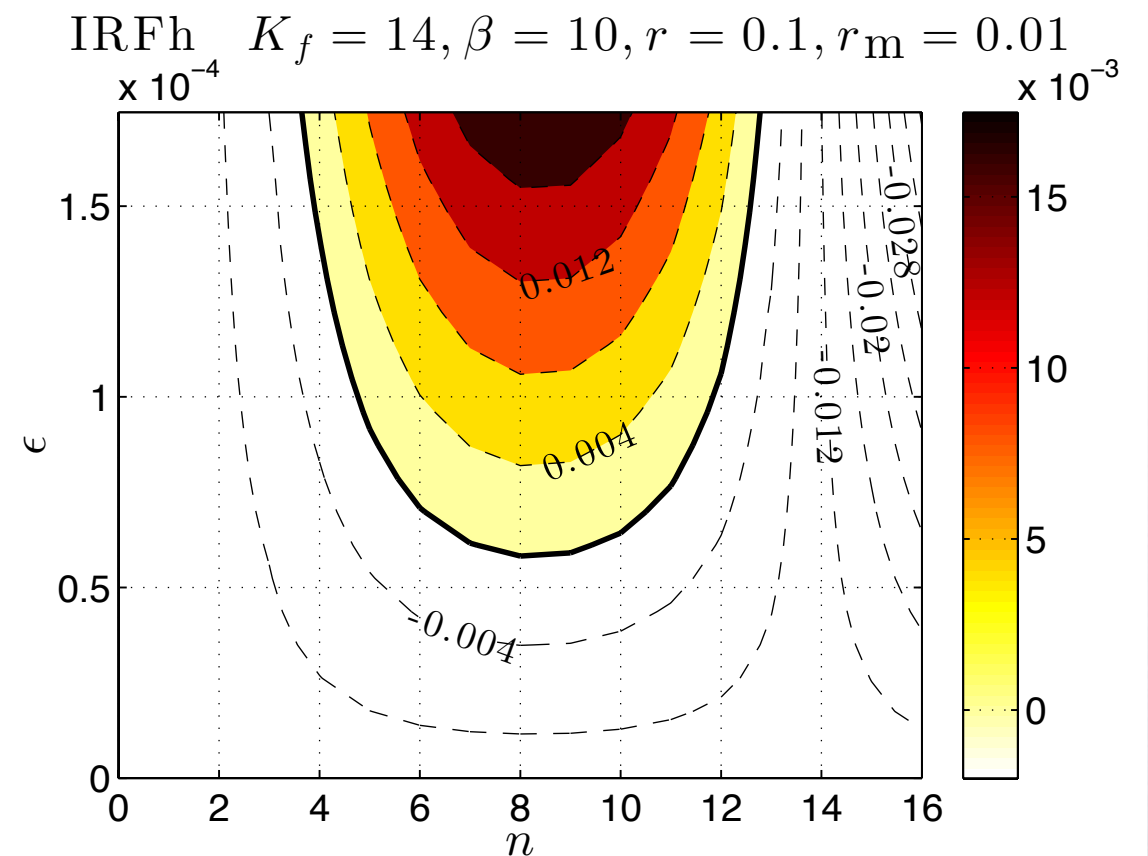
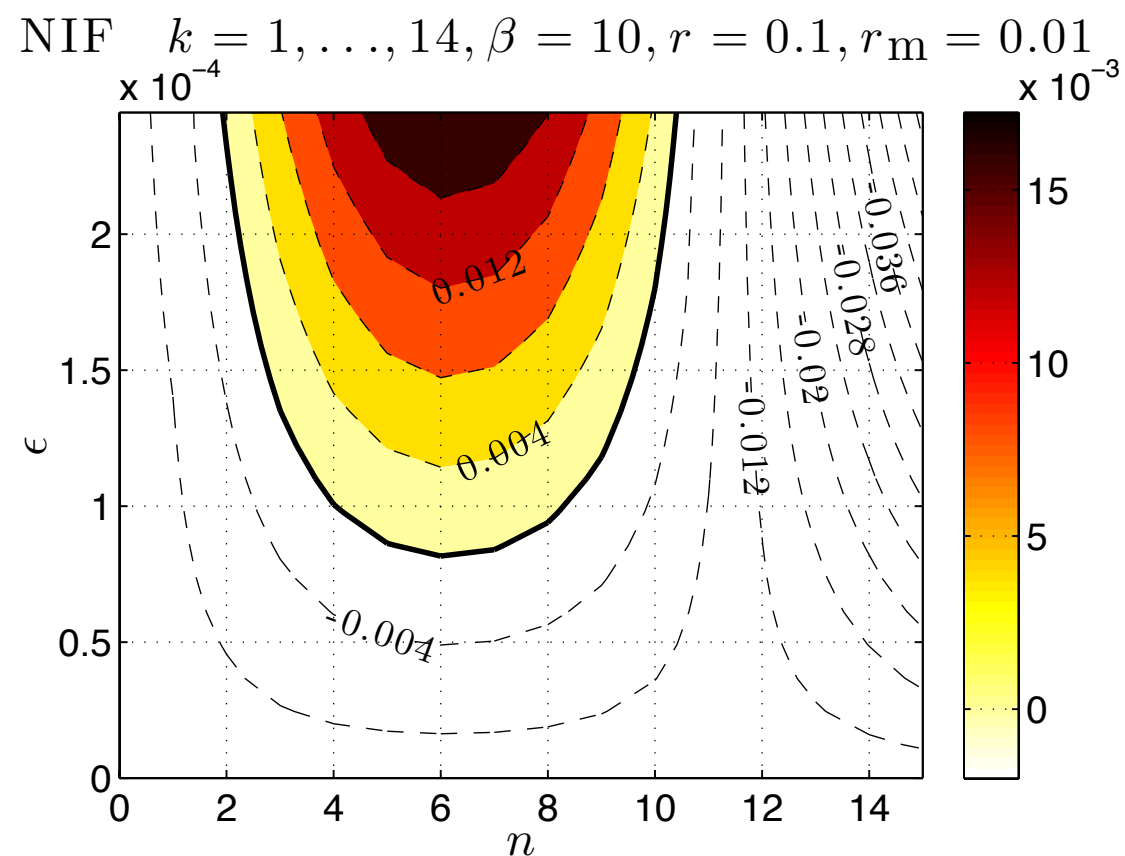
Isotropic Ring Forcing (IRFh)

parametrize excitation by
convection





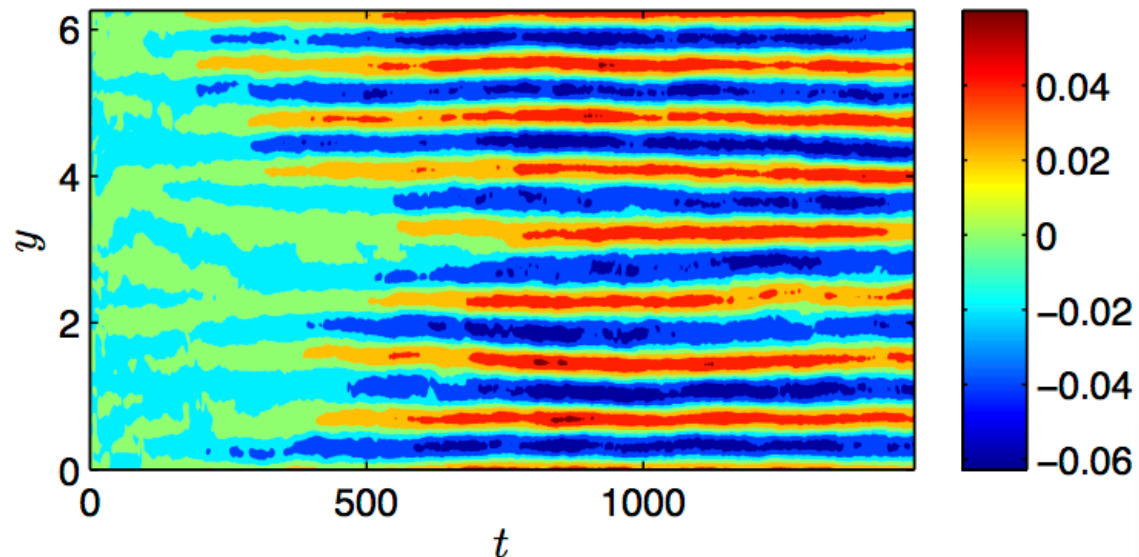
SSST growth rates for homogeneous equilibrium



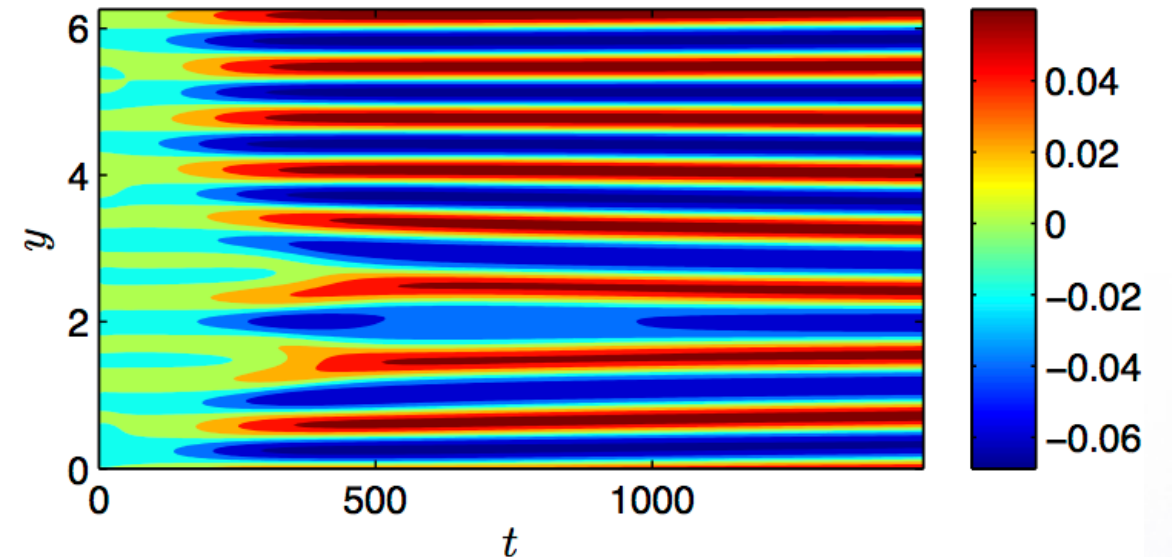


IRFh forcing, $r_m = r/10$ (NL, QL, SSST)

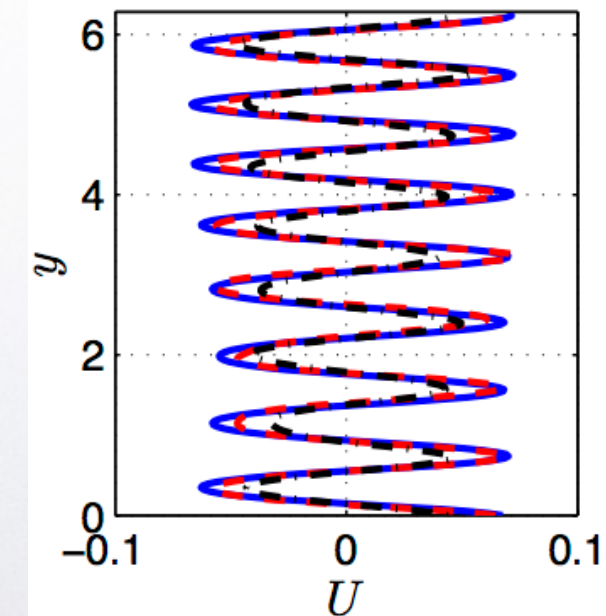
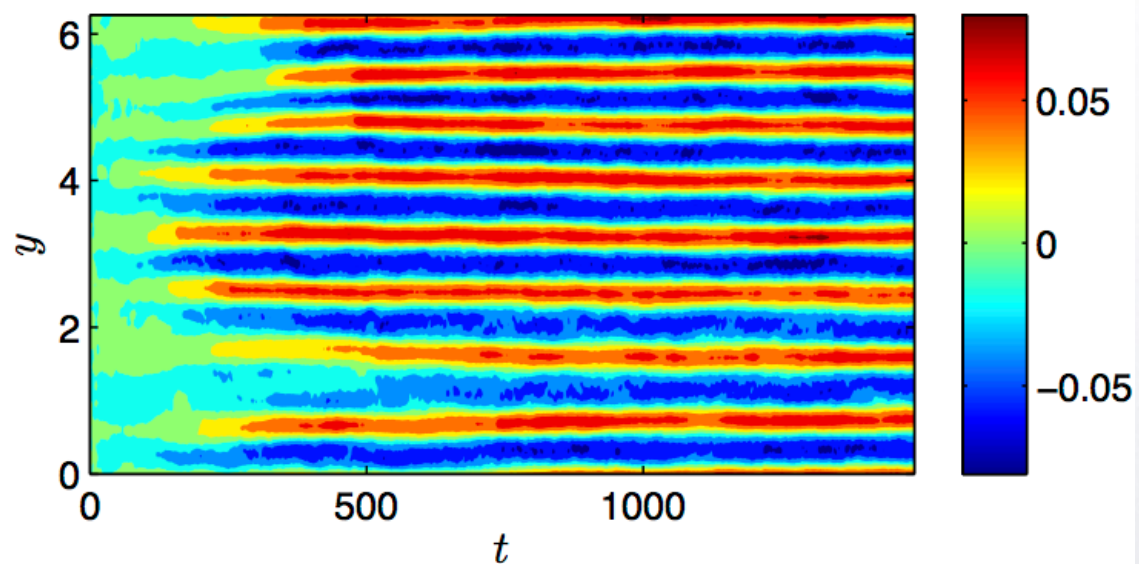
NL $U(y, t)$, $\epsilon/\epsilon_c = 2.5$



SSST $U(y, t)$, $\epsilon/\epsilon_c = 2.5$



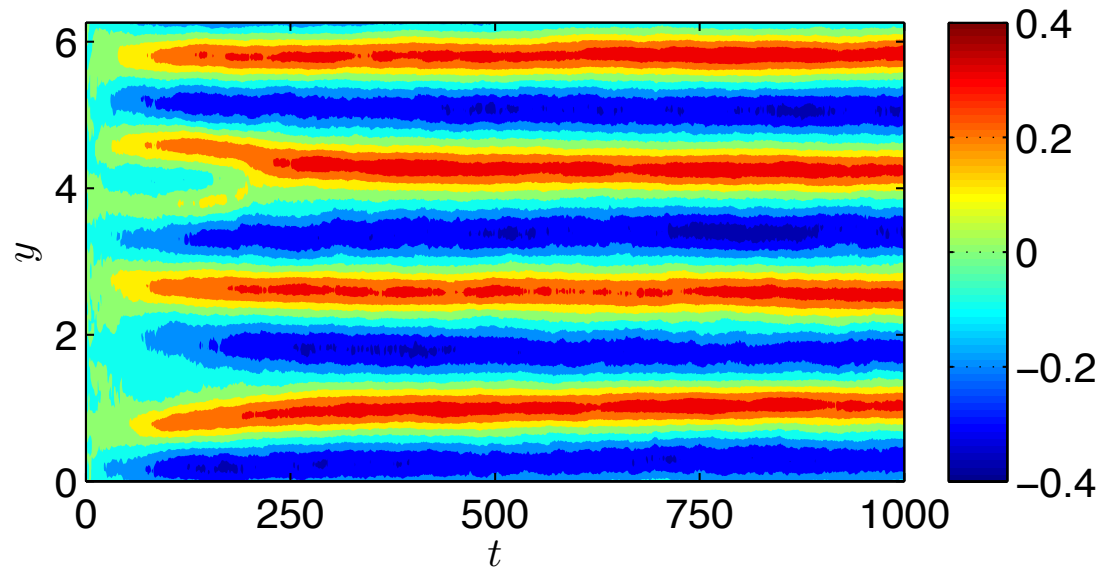
QL $U(y, t)$, $\epsilon/\epsilon_c = 2.5$



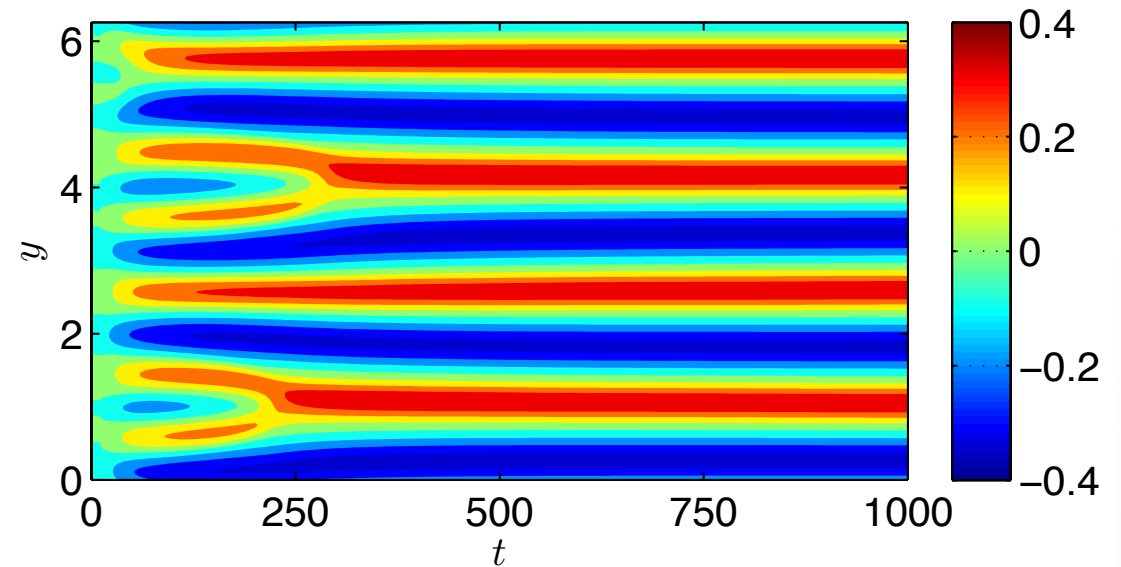


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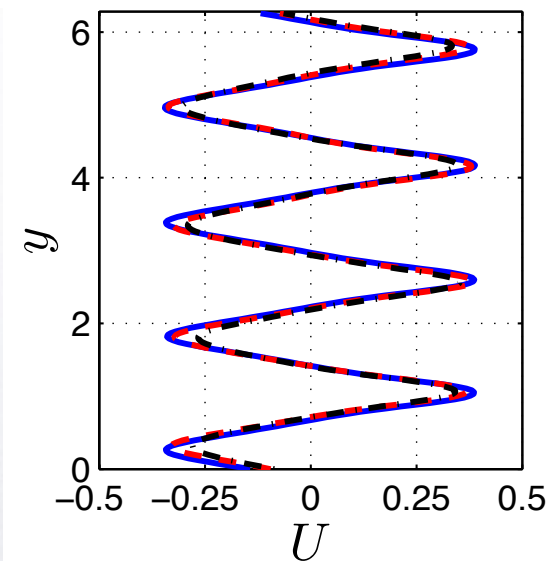
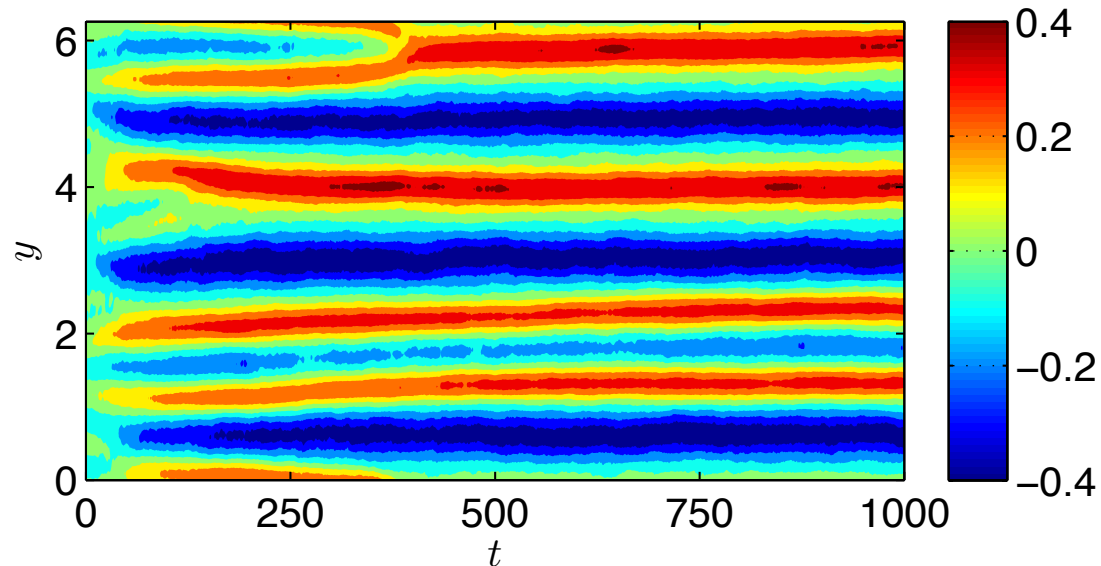
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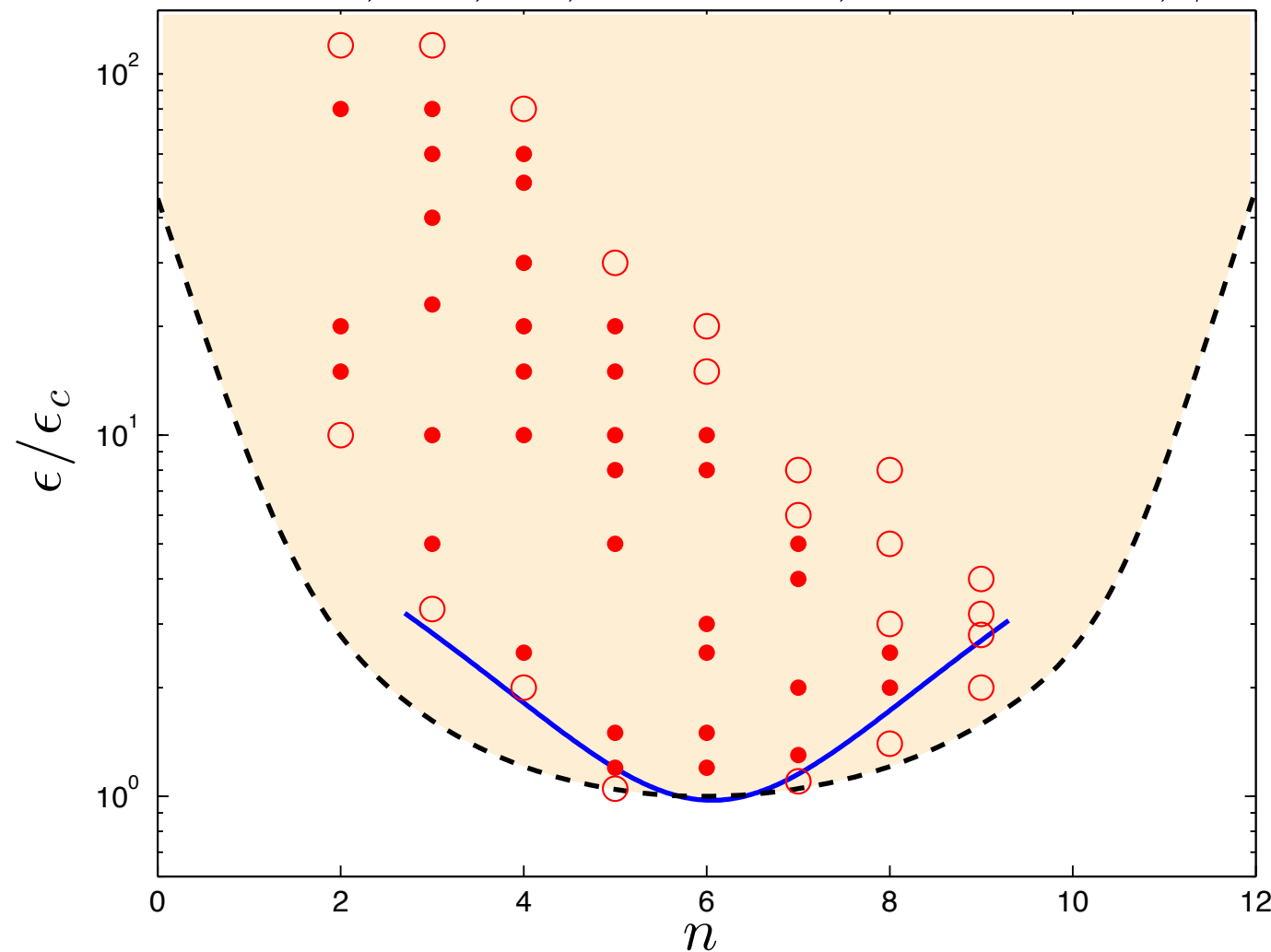
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SSST equilibria stability diagram

NIF at $k = 1, \dots, 14$, $r = 10^{-1}$, $r_m = 10^{-2}$, $\beta = 10$

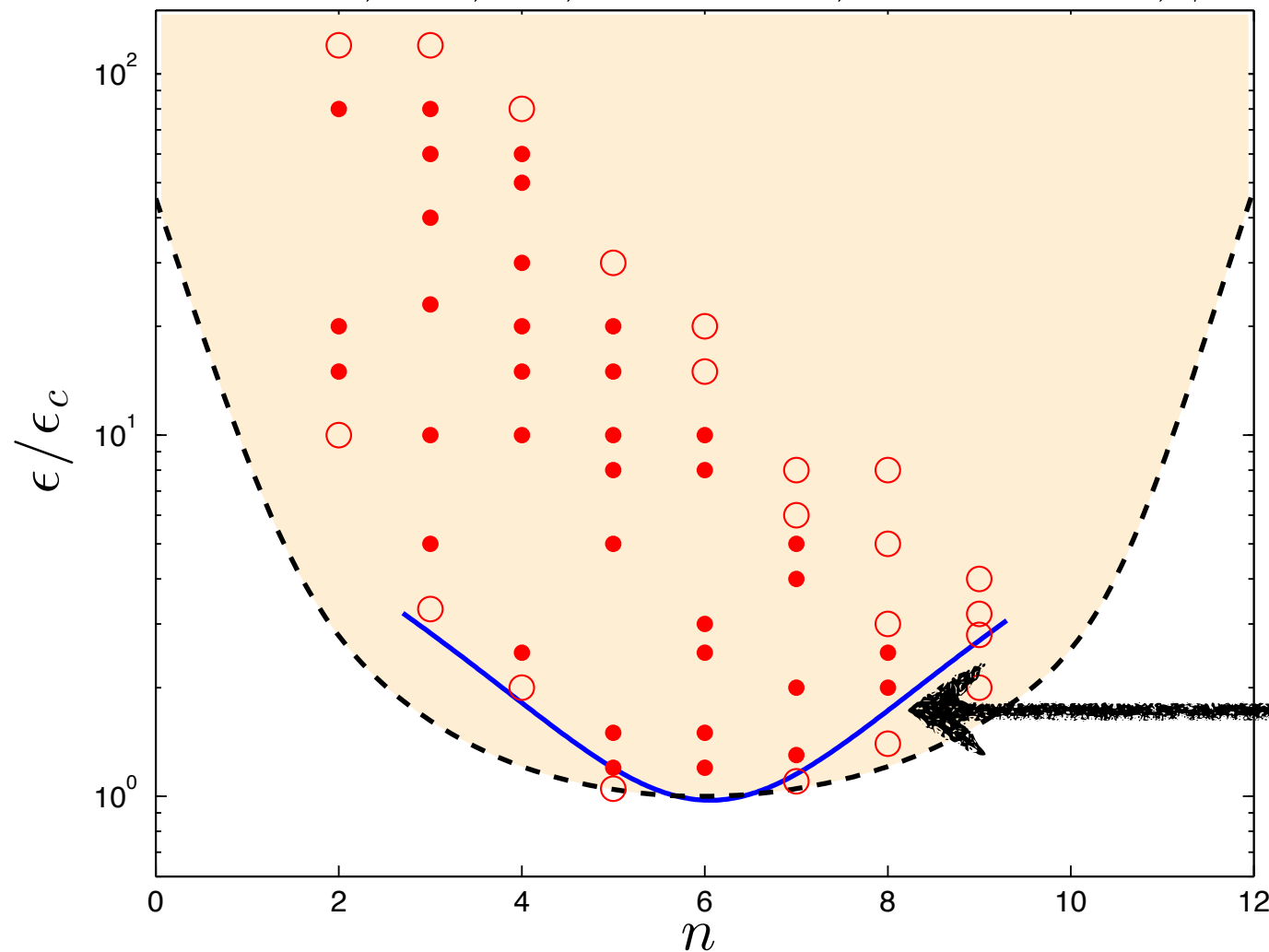


For higher energy input rates equilibria become SSST unstable and move towards the left of the diagram



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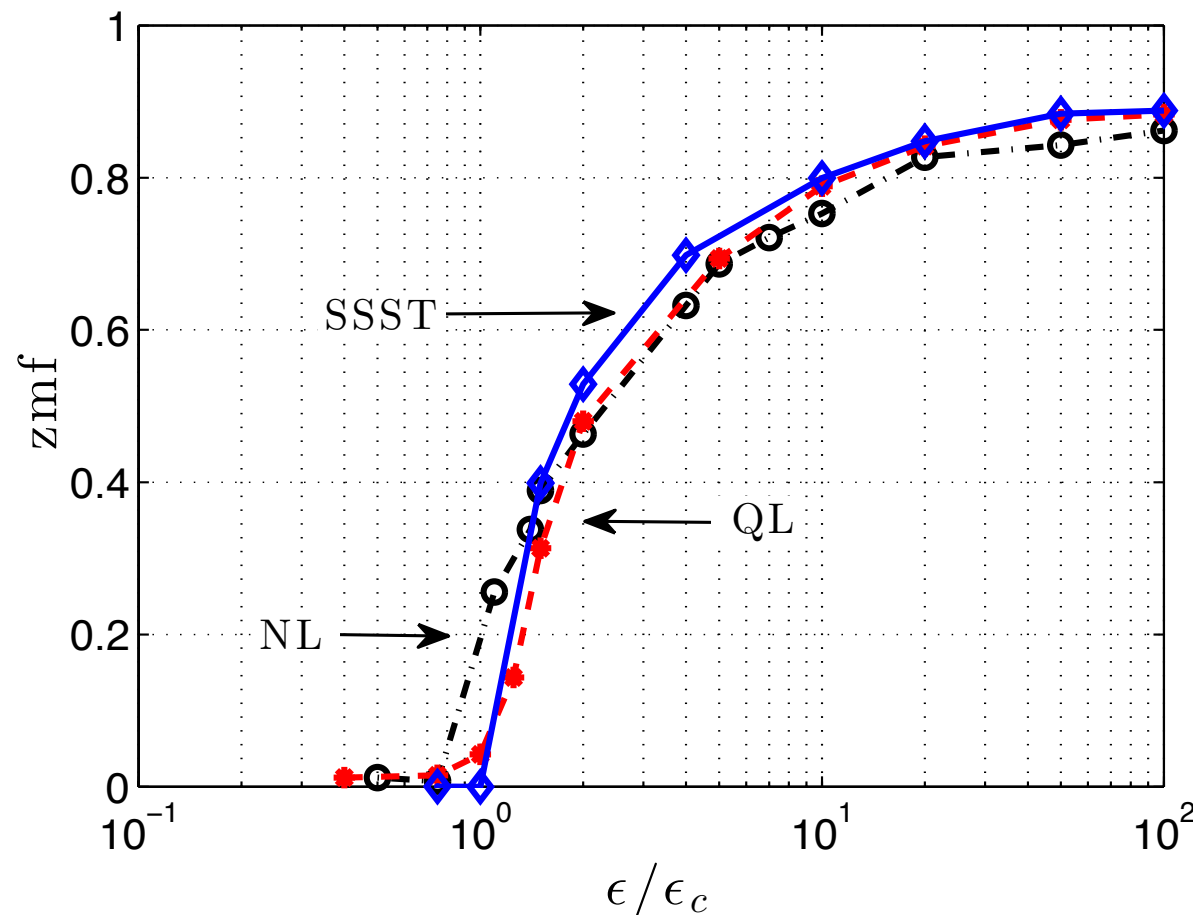
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Eckhaus instability boundary

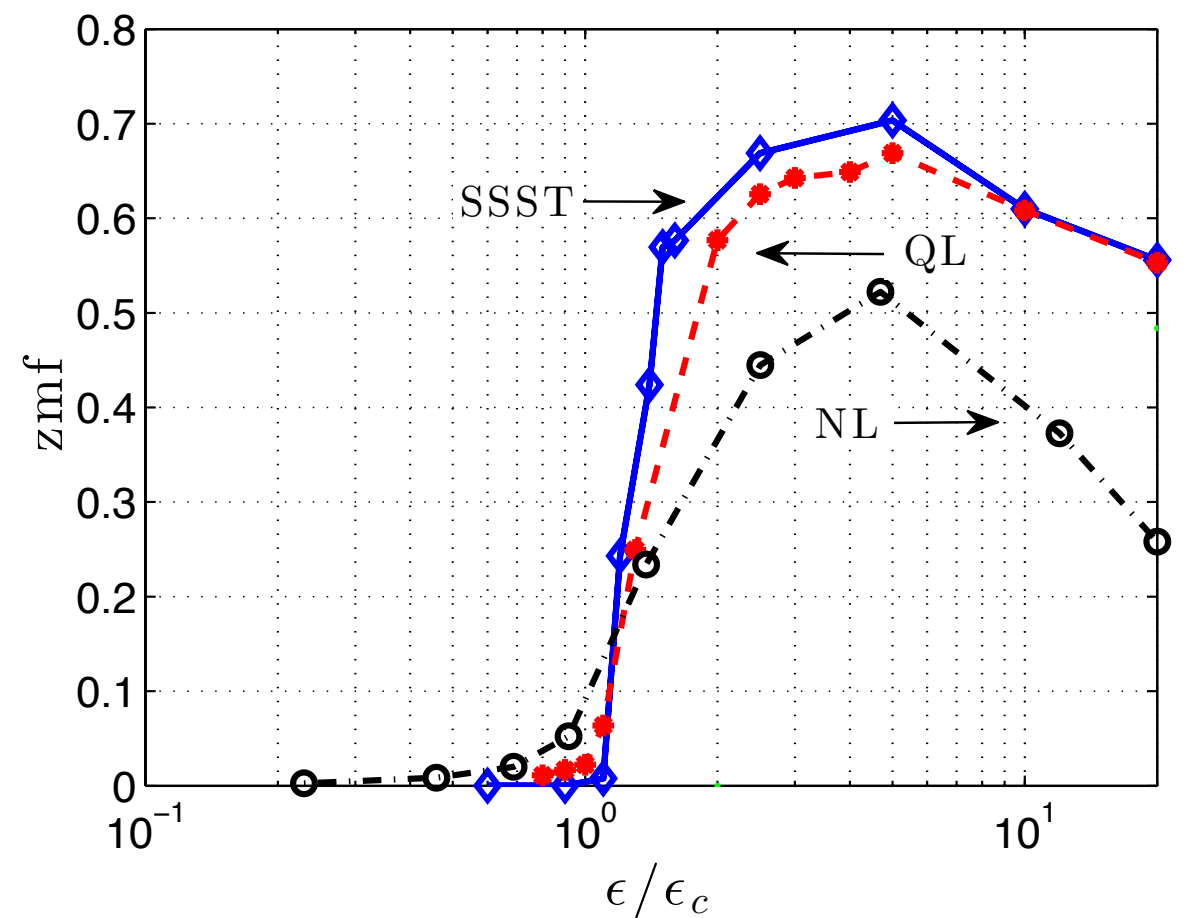


Bifurcation diagrams

NIF, $r = 0.1$, $r_m = r/10$



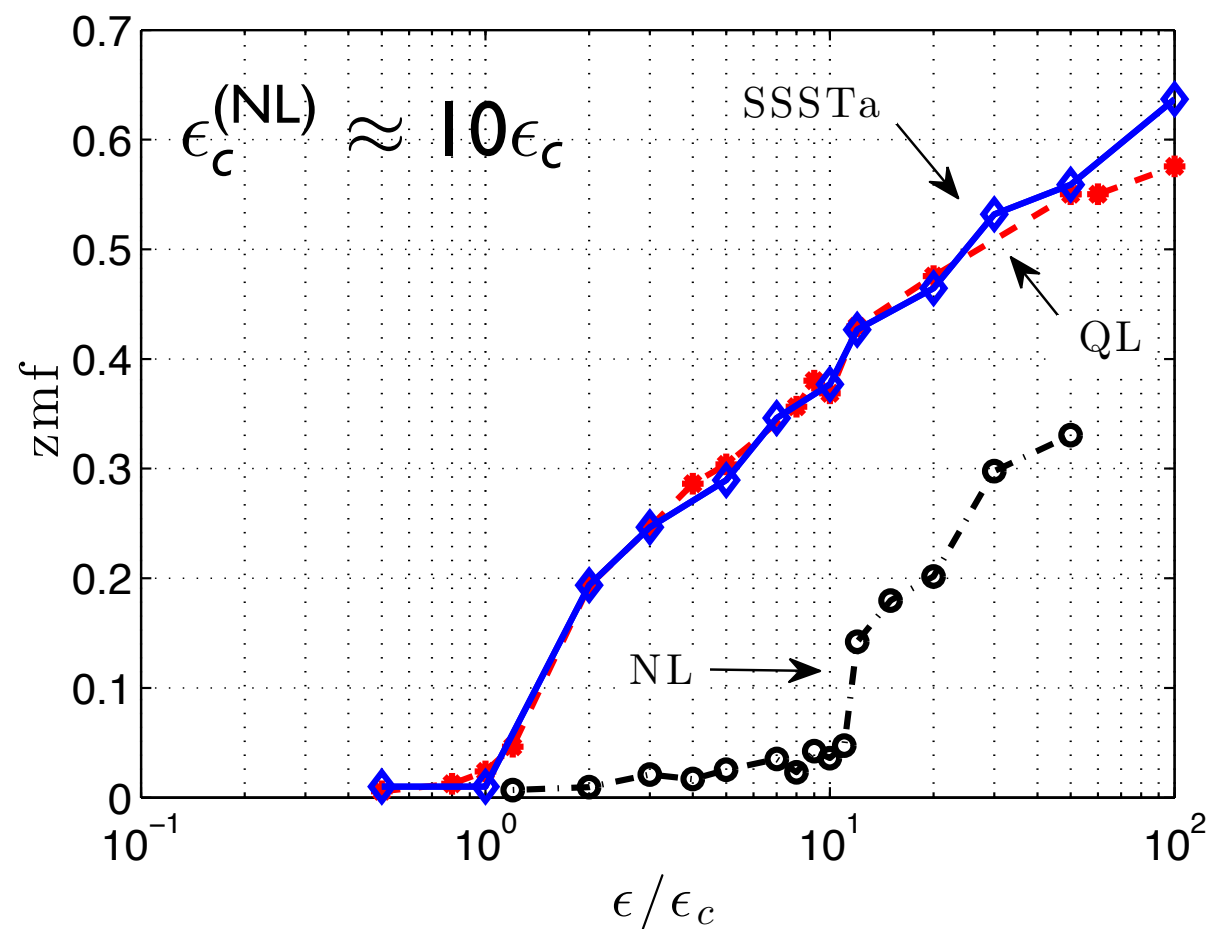
IRFh, $r = 0.1$, $r_m = r/10$



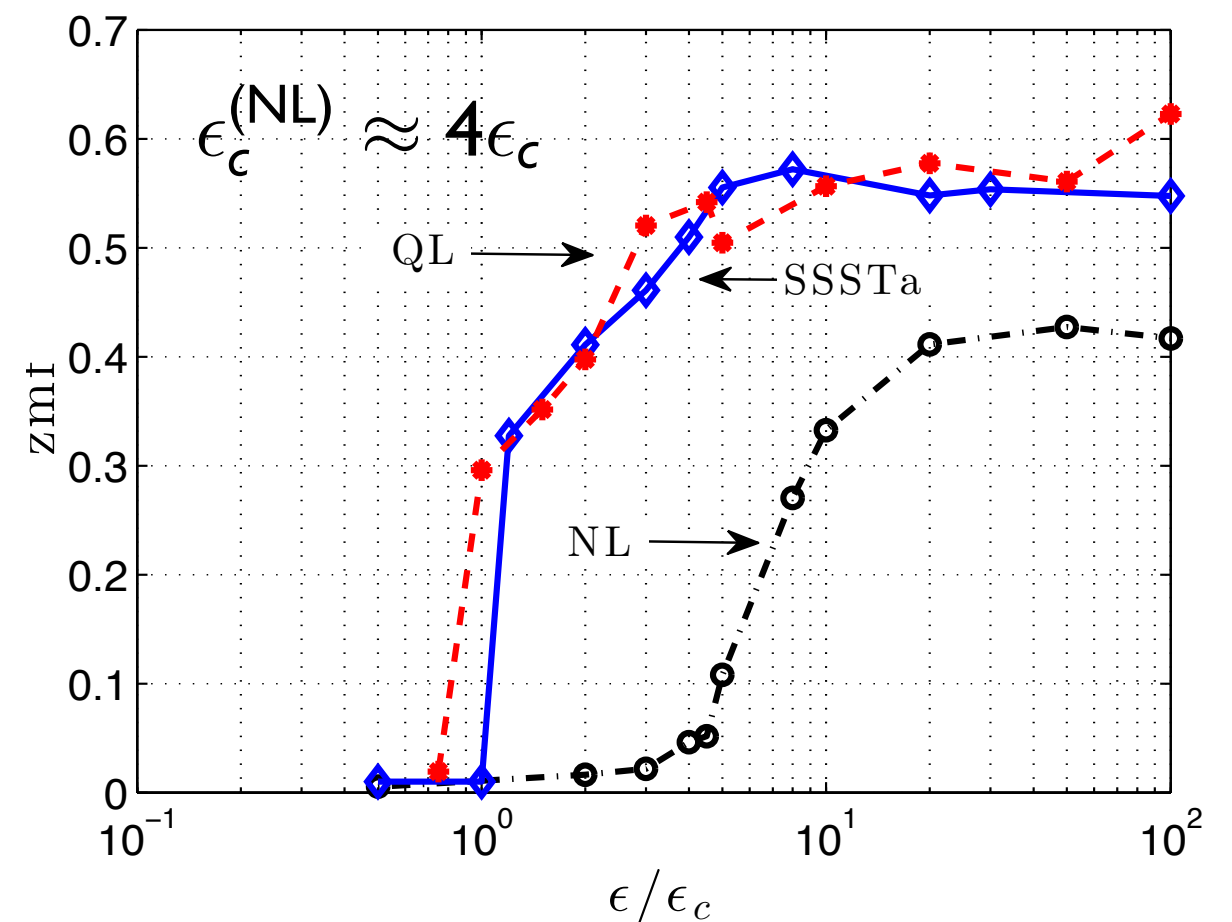


Bifurcation diagrams

NIF, $r = r_m = 10^{-2}$



IRFh, $r = r_m = 10^{-2}$

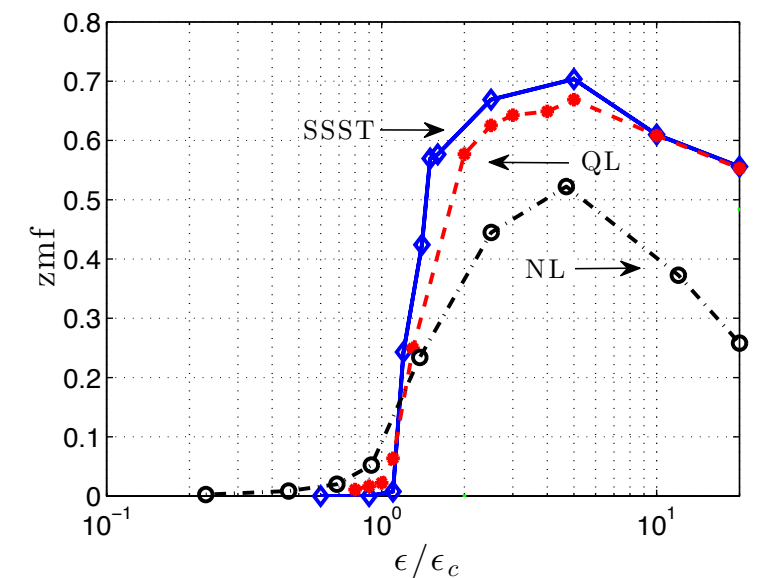




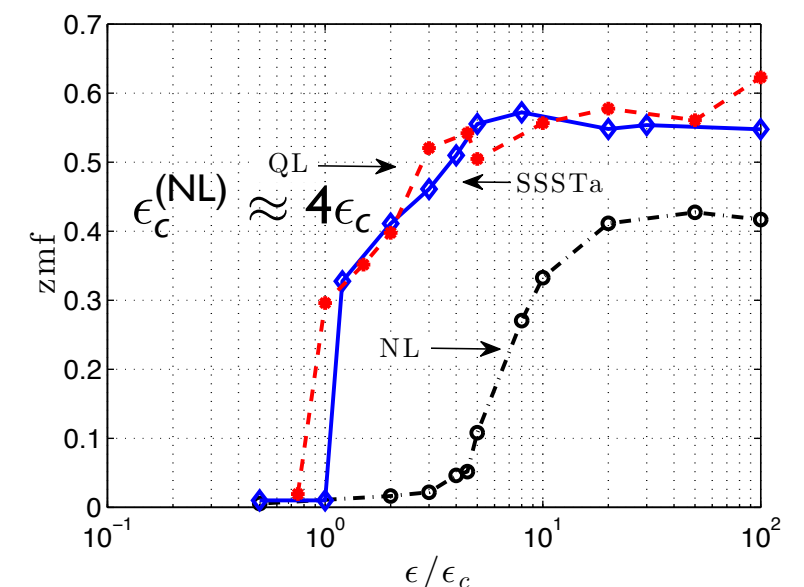
Issues to be clarified

- ▶ What happens below the critical energy input rate, ϵ_c ?
- ▶ When does the predicted ϵ_c agrees with the NL bifurcation point, $\epsilon_c^{(NL)}$ and why they don't always agree?
- ▶ What happens in the range $\epsilon_c < \epsilon < \epsilon_c^{(NL)}$?
- ▶ How can $\epsilon_c^{(NL)}$ be predicted?

IRFh, $r = 0.1$, $r_m = r/10$



IRFh, $r = r_m = 10^{-2}$

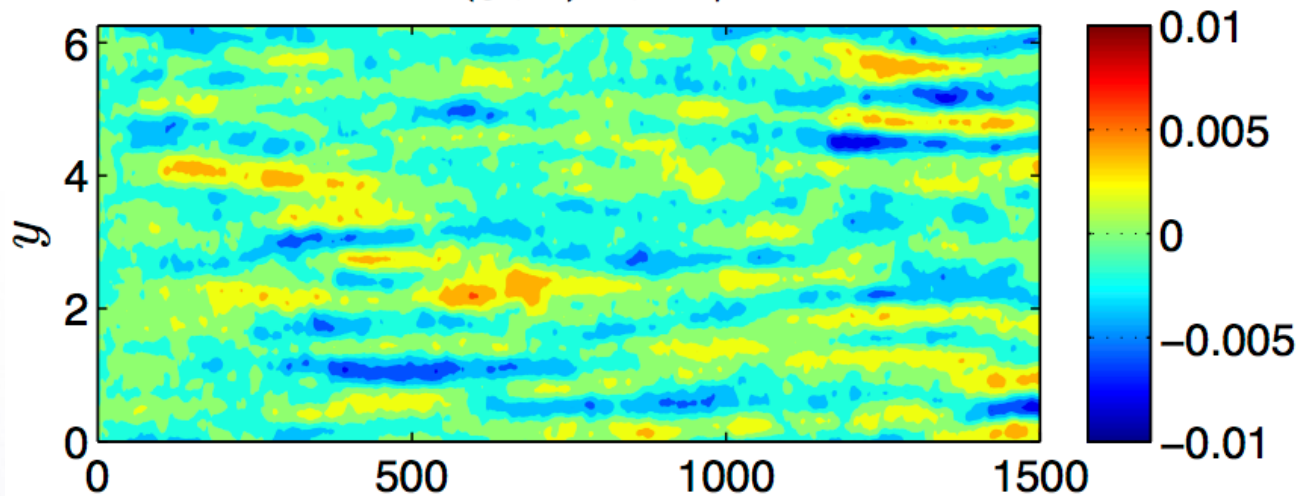




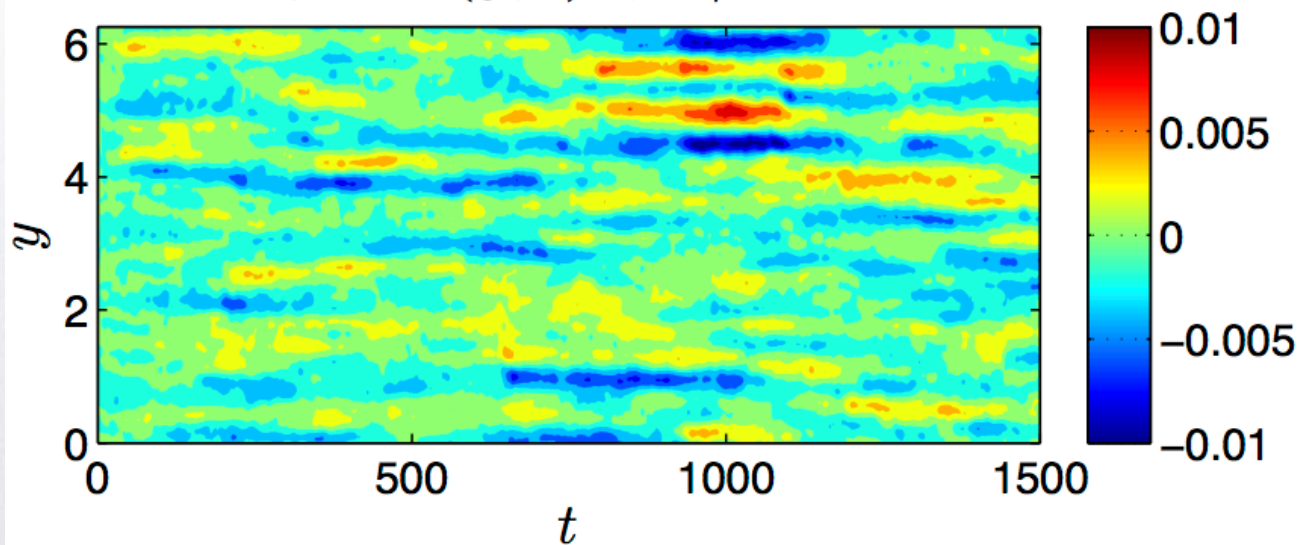
Stochastically excited stable SSST eigenmodes

(Latent jets in ocean: Berloff et. al. 2009, Berloff et. al. 2011, Cravatte et. al. 2012)

NL $U(y, t)$, $\epsilon/\epsilon_c = 0.8$



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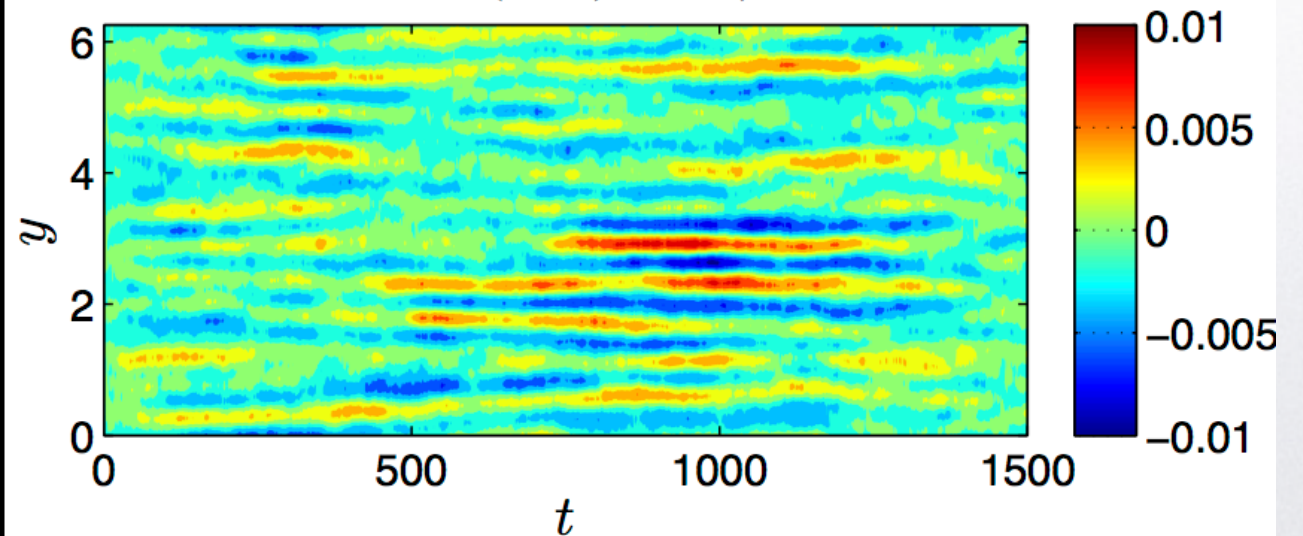
$$U(y, t) = \text{Re} \left[\sum_{n=1}^{15} \alpha_n(t) e^{iny} \right]$$

$$\frac{d\alpha_n}{dt} = \sigma_n \alpha_n + \xi(t)$$

σ_n : SSST growth rates (least stable)

$\xi(t)$: delta-correlated noise

SSST $U(y, t)$, $\epsilon/\epsilon_c = 0.8$

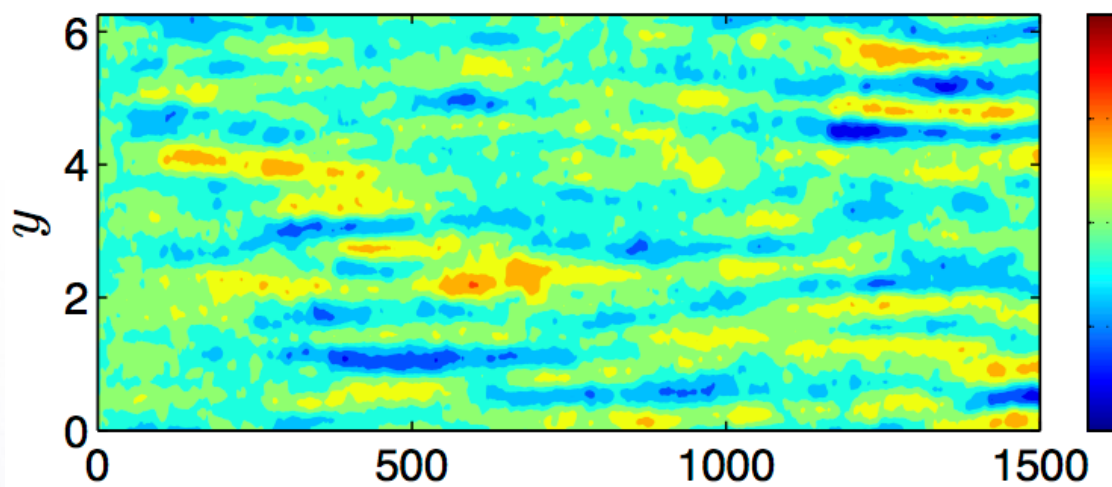




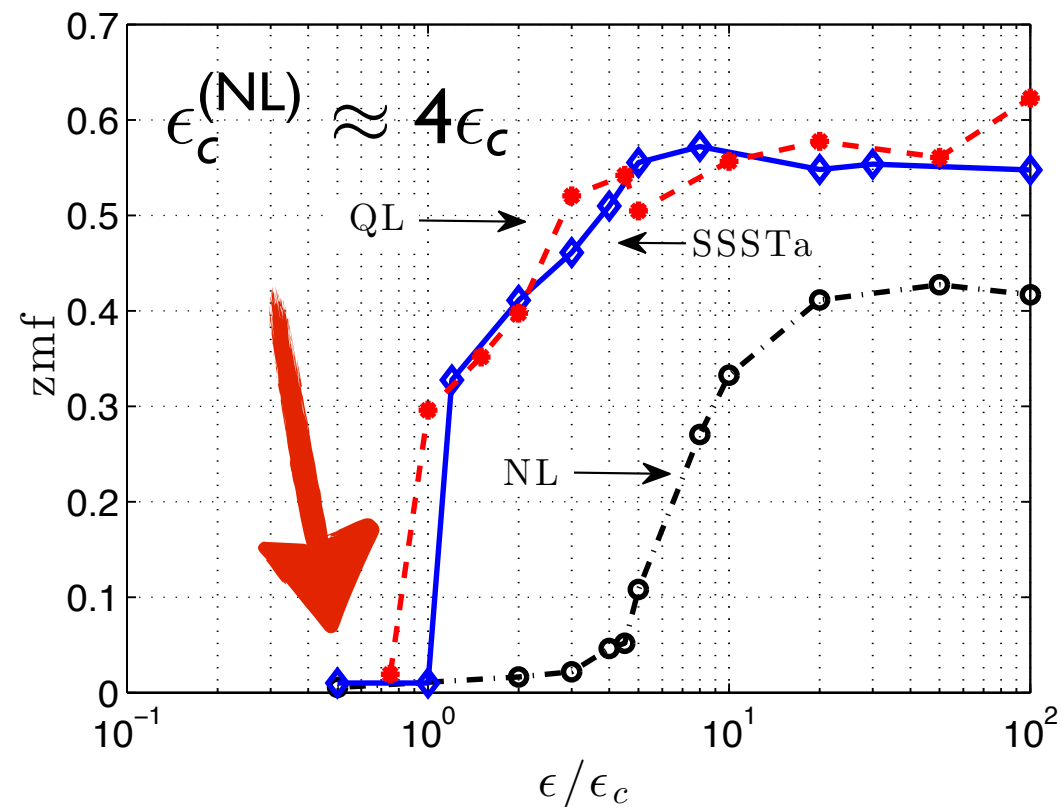
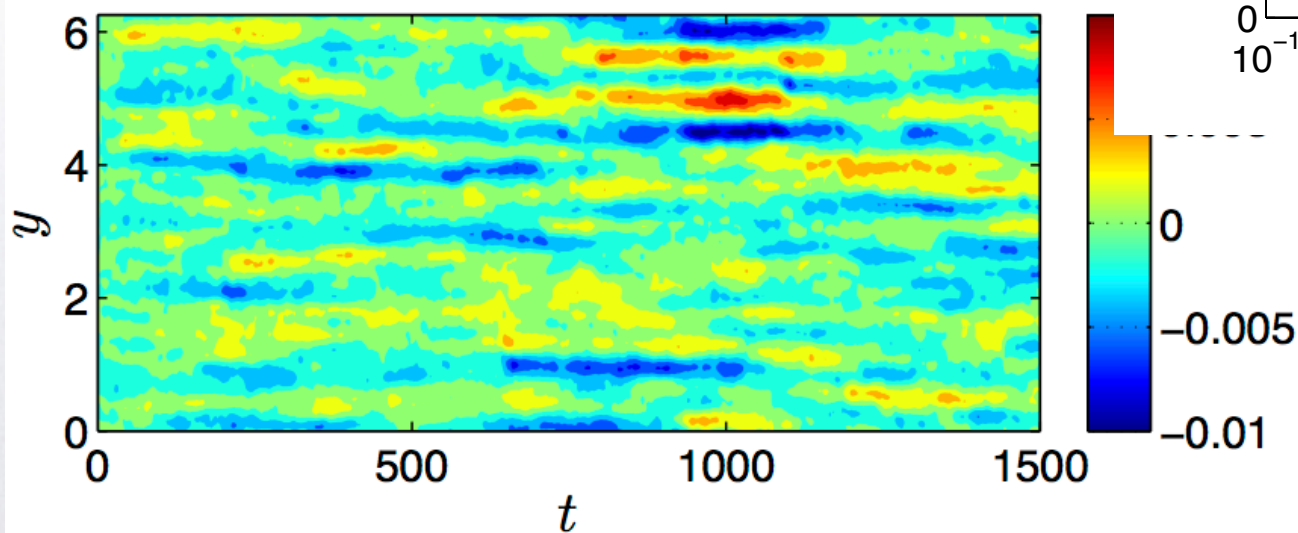
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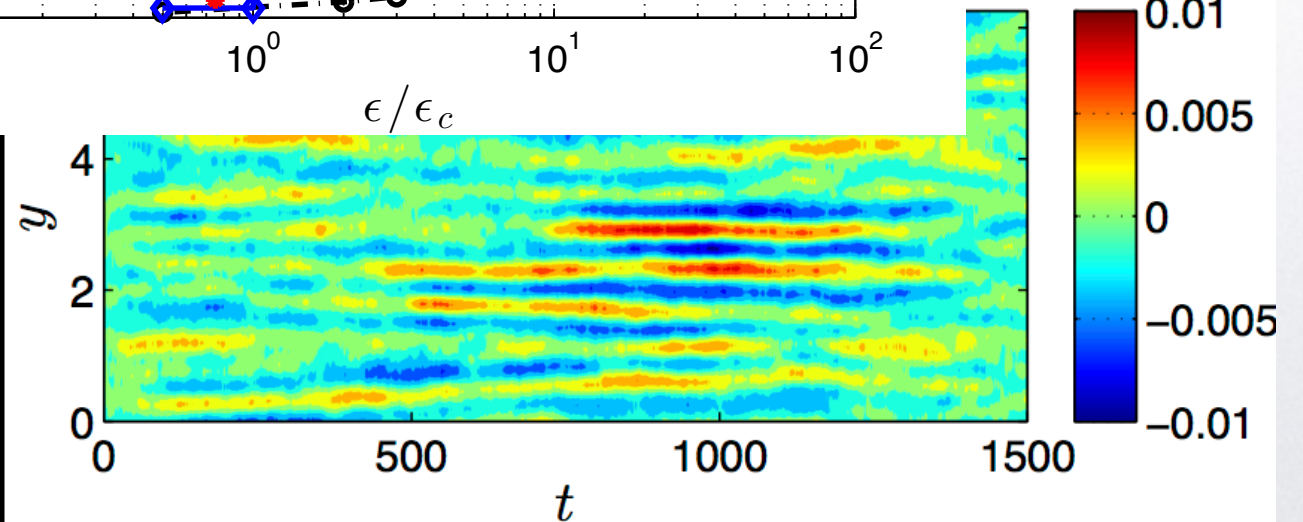
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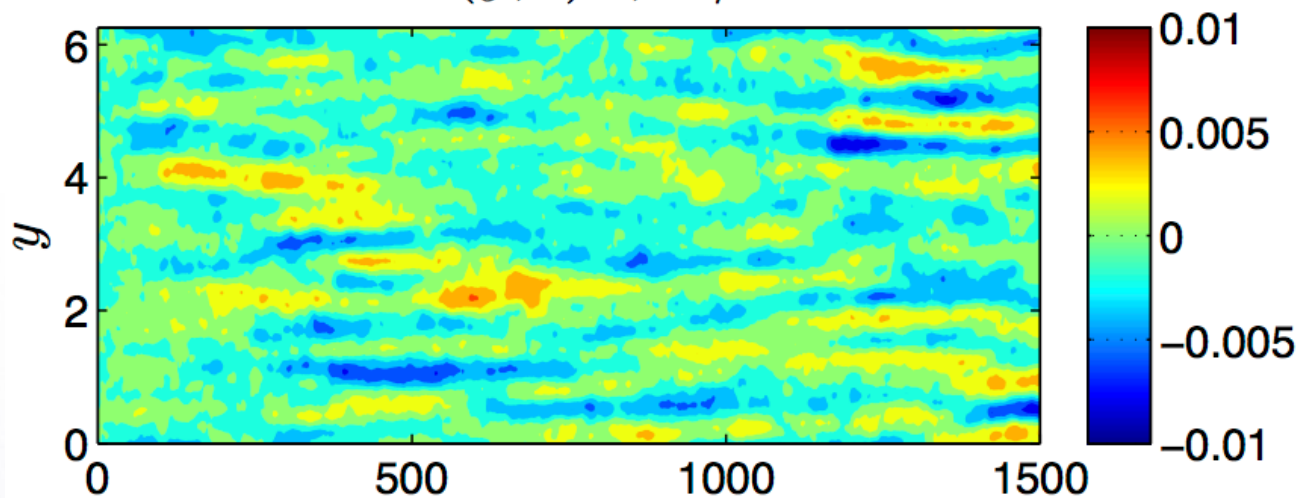




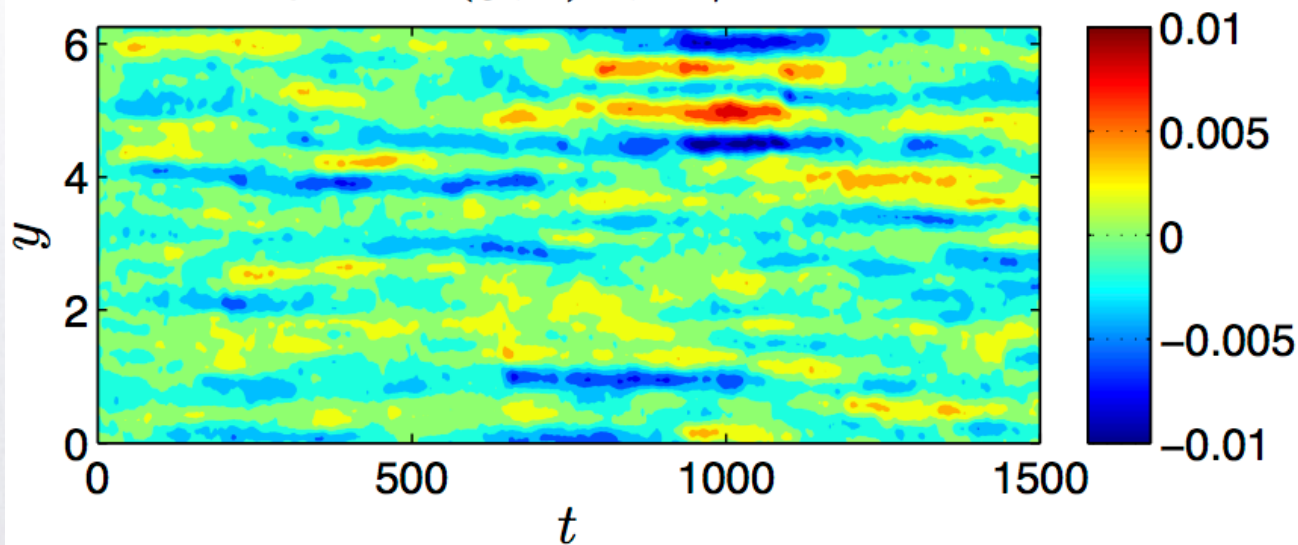
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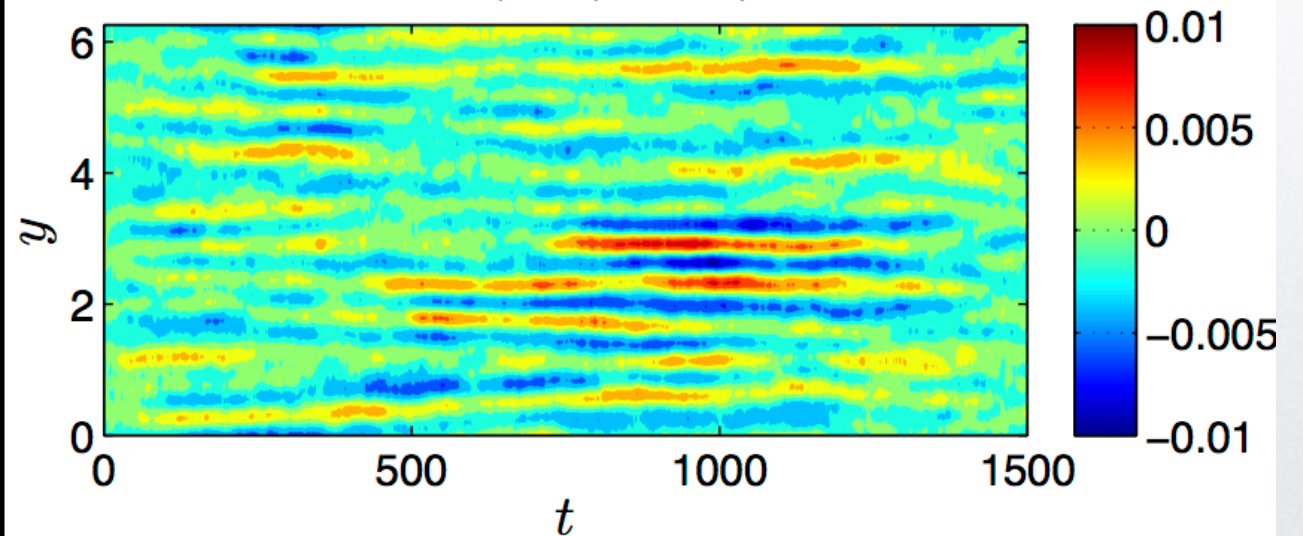
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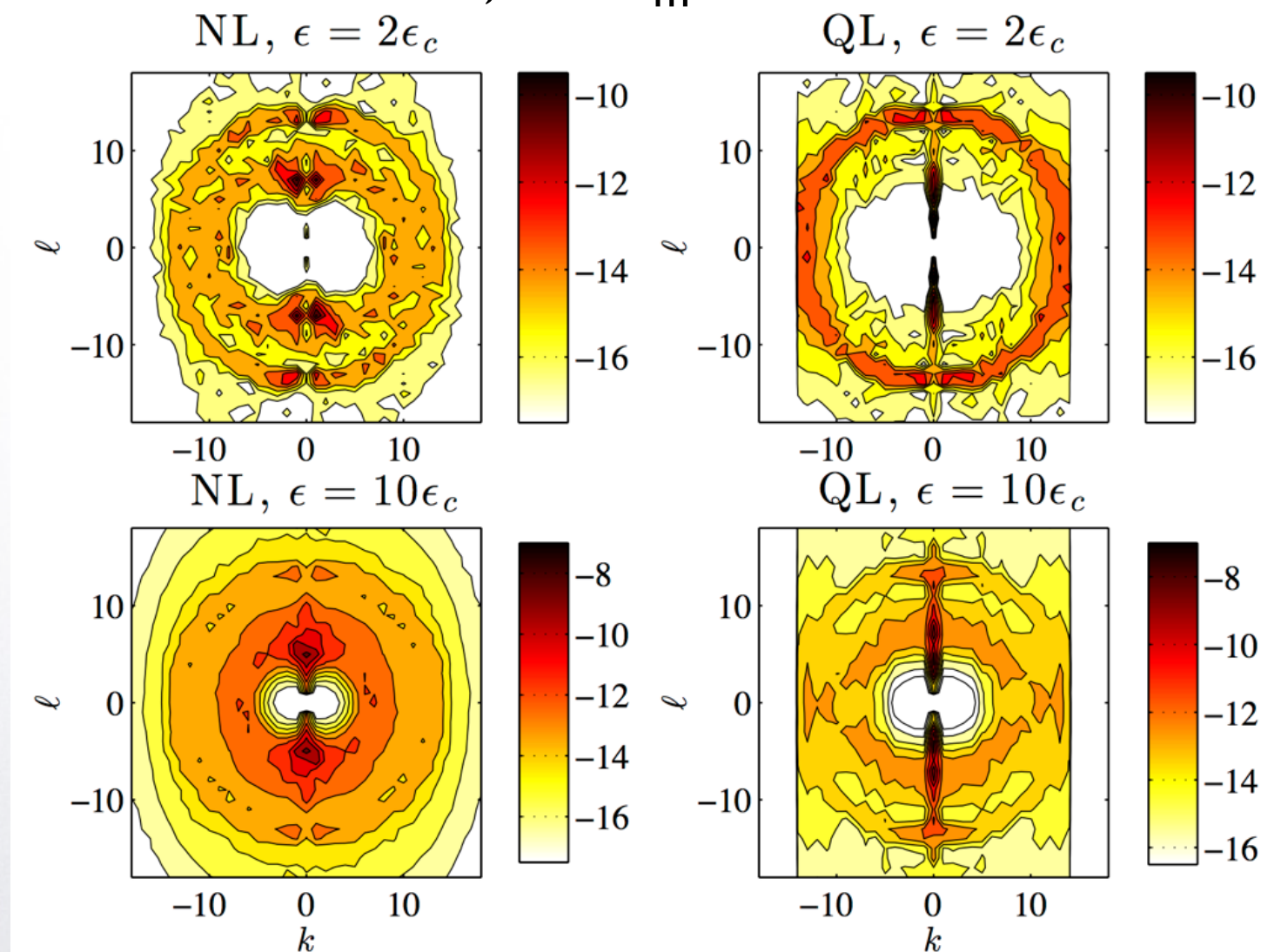




Zonons at NL

energy spectrum

IRFh, $r = r_m = 10^{-2}$



QL / SSST do not allow the the interaction of the turbulence with coherent non-zonal structures



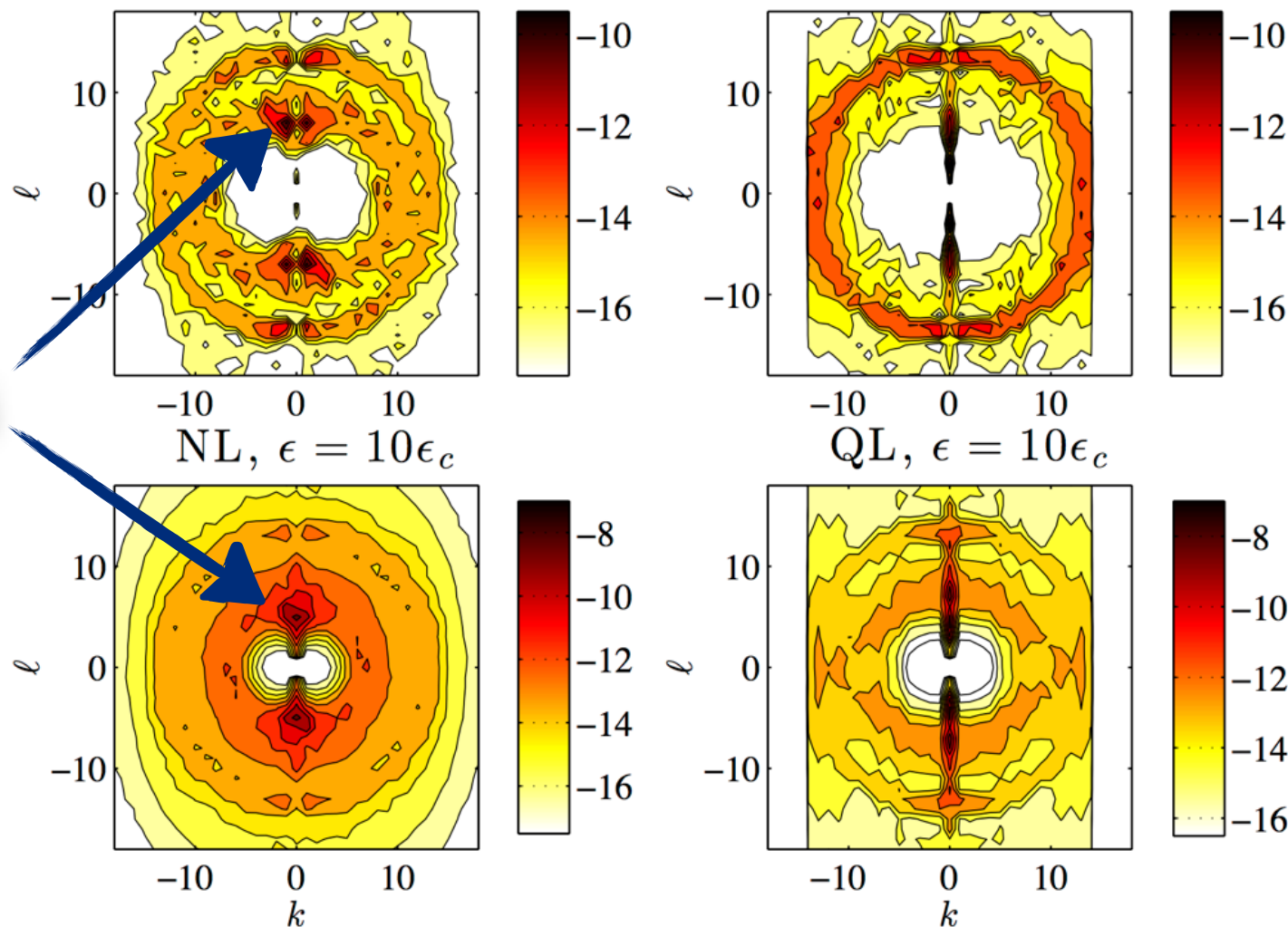
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Zonons

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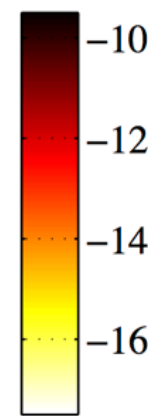
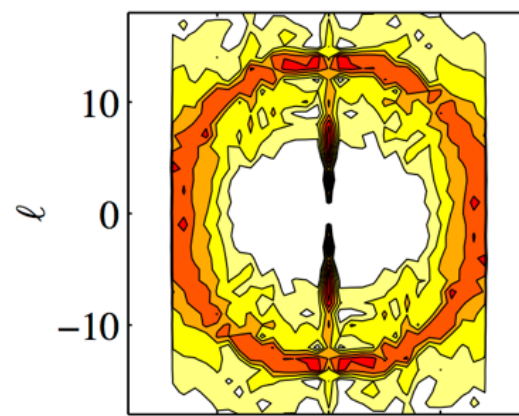
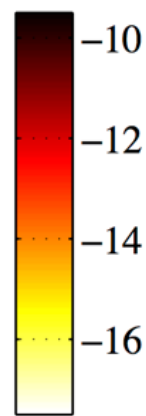
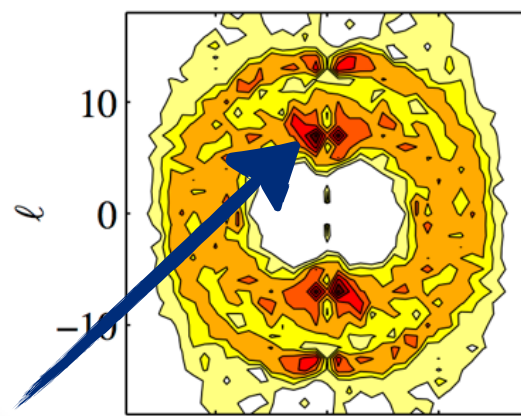
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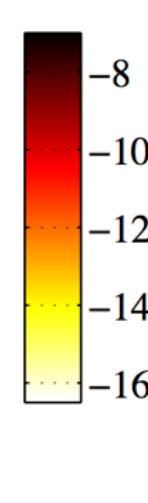
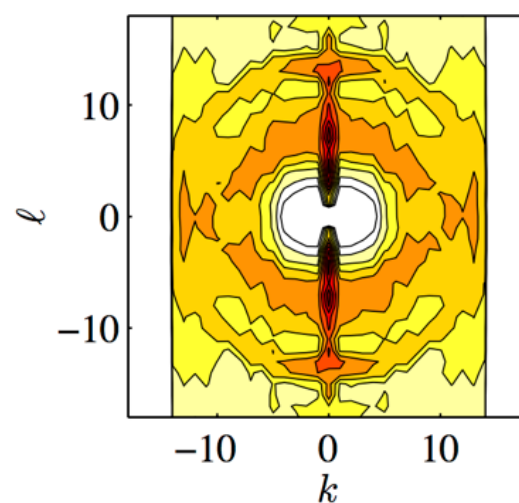
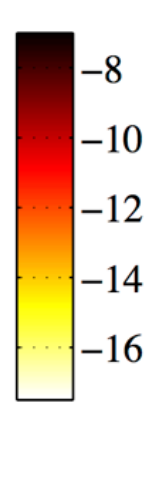
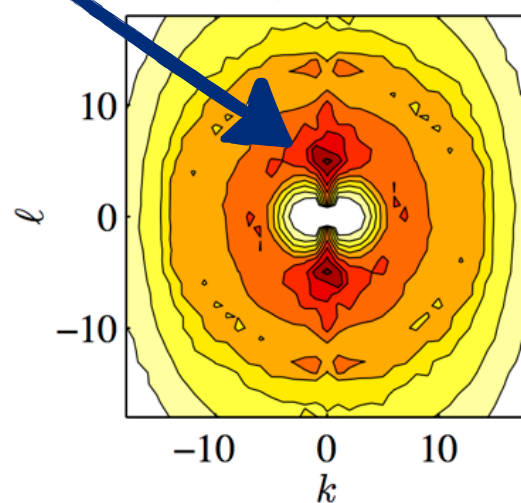
NL, $\epsilon = 2\epsilon_c$

QL, $\epsilon = 2\epsilon_c$



NL, $\epsilon = 10\epsilon_c$

QL, $\epsilon = 10\epsilon_c$



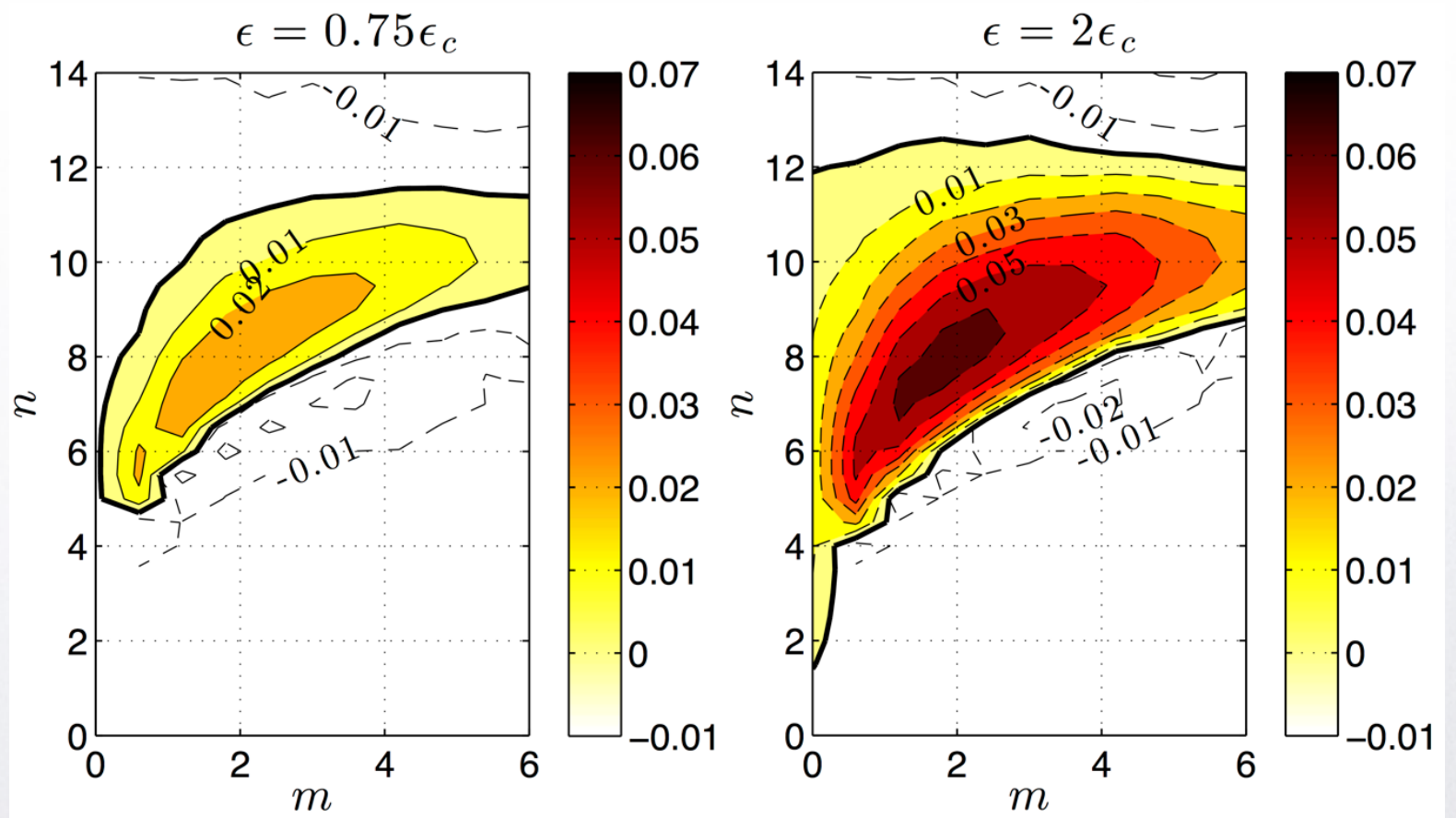
Zonons

QL / SSST do not allow the the interaction of the turbulence with coherent non-zonal structures



perturbations of the
homogeneous equilibrium in
form $\delta U = e^{i(mx+ny)}$

for these parameters non-zonal perturbations are more SSST unstable than jets ($m = 0$)

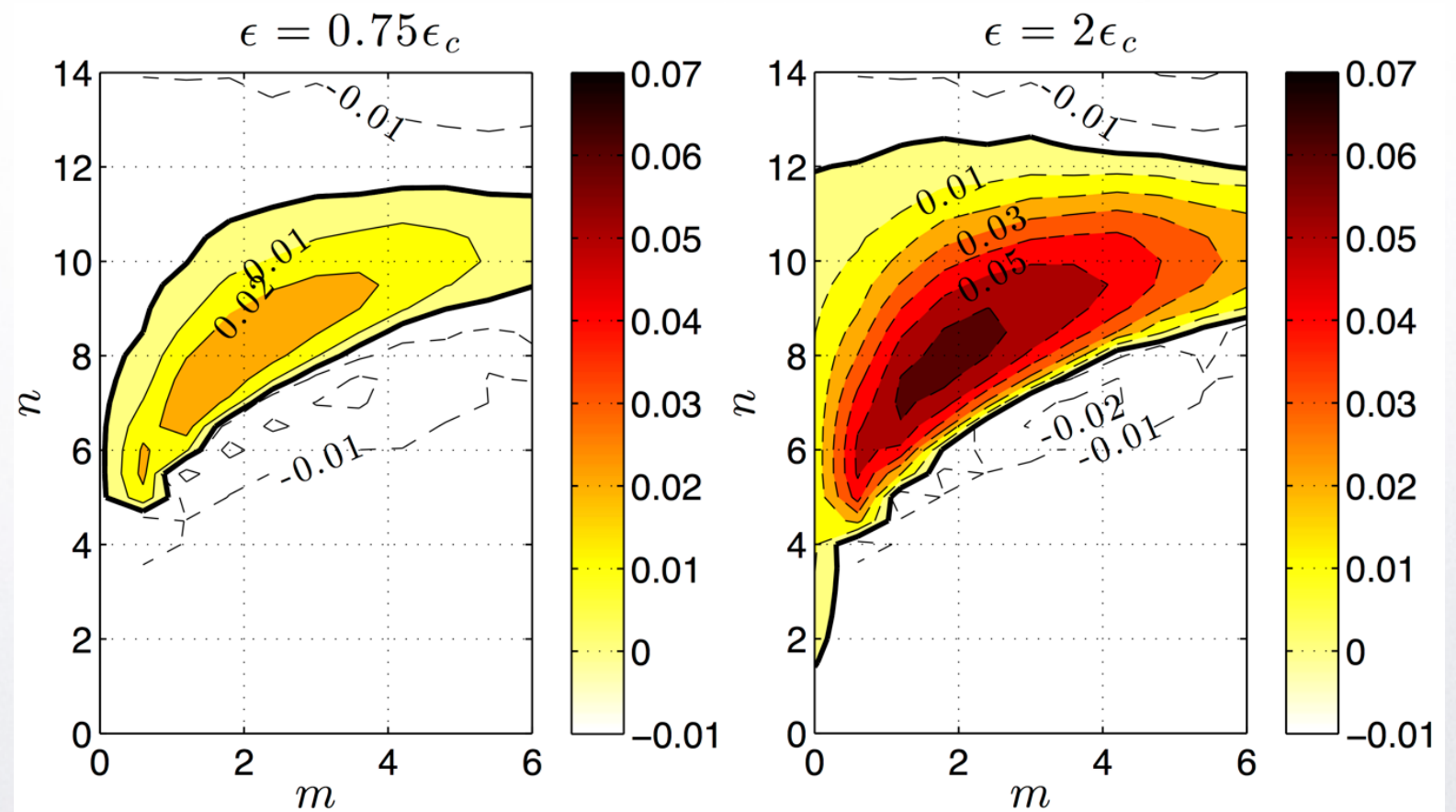
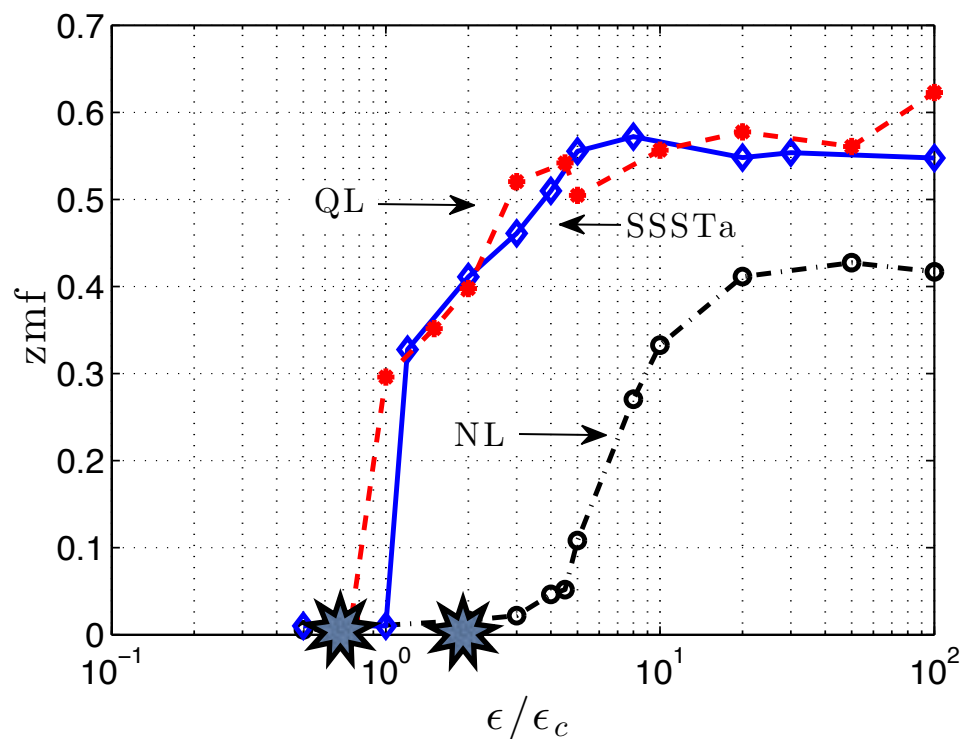




Non-zonal instabilities to homogeneous equilibrium

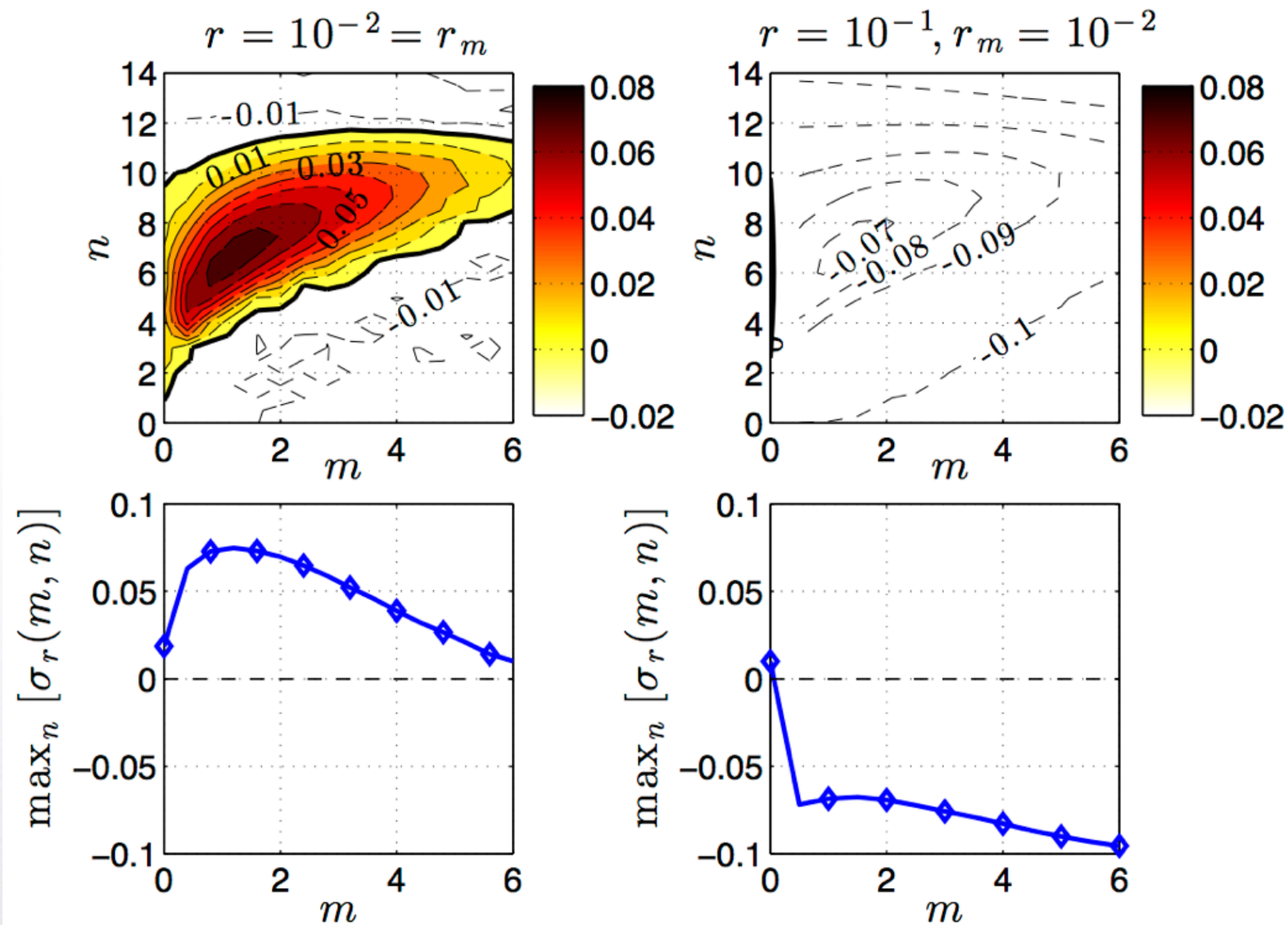
IRFh, $r = r_m = 10^{-2}$

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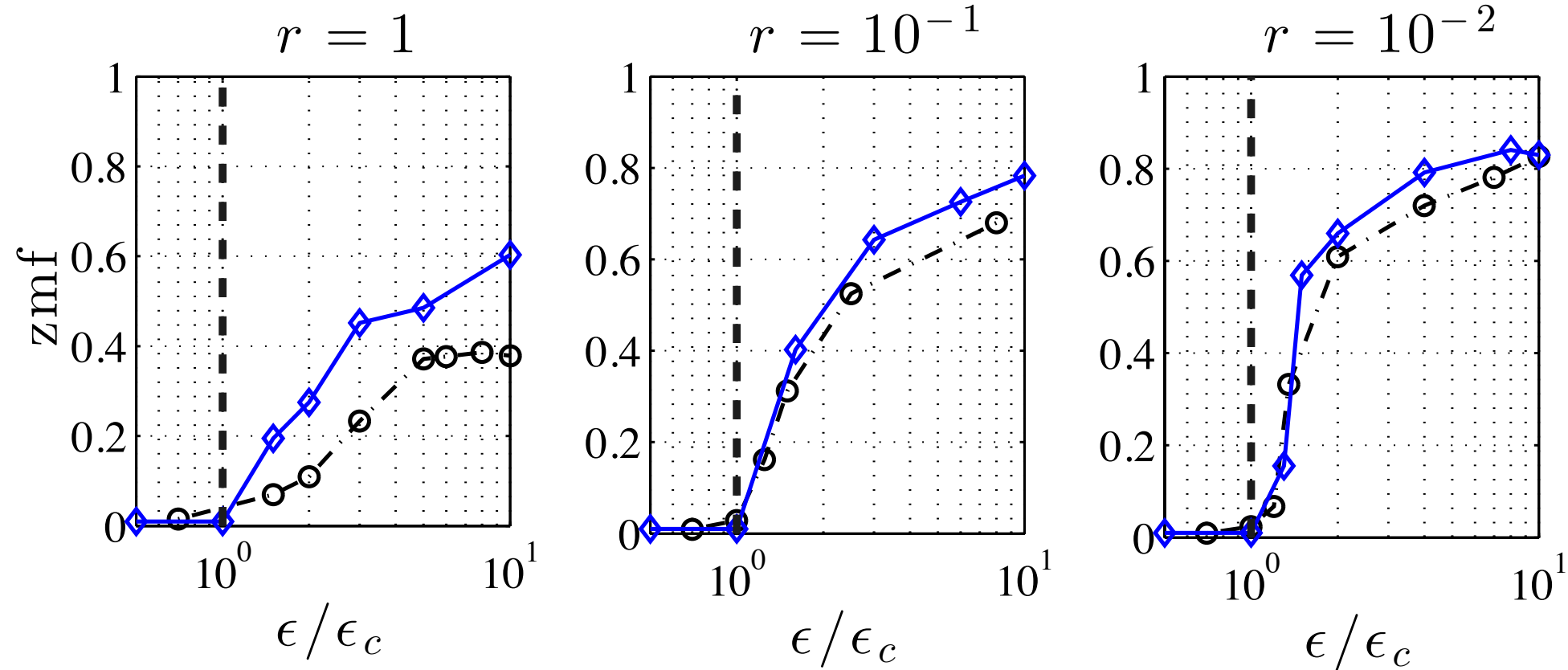
Depression of non-zonal instabilities using $r_m = r/10$





SSST captures the NL bifurcation

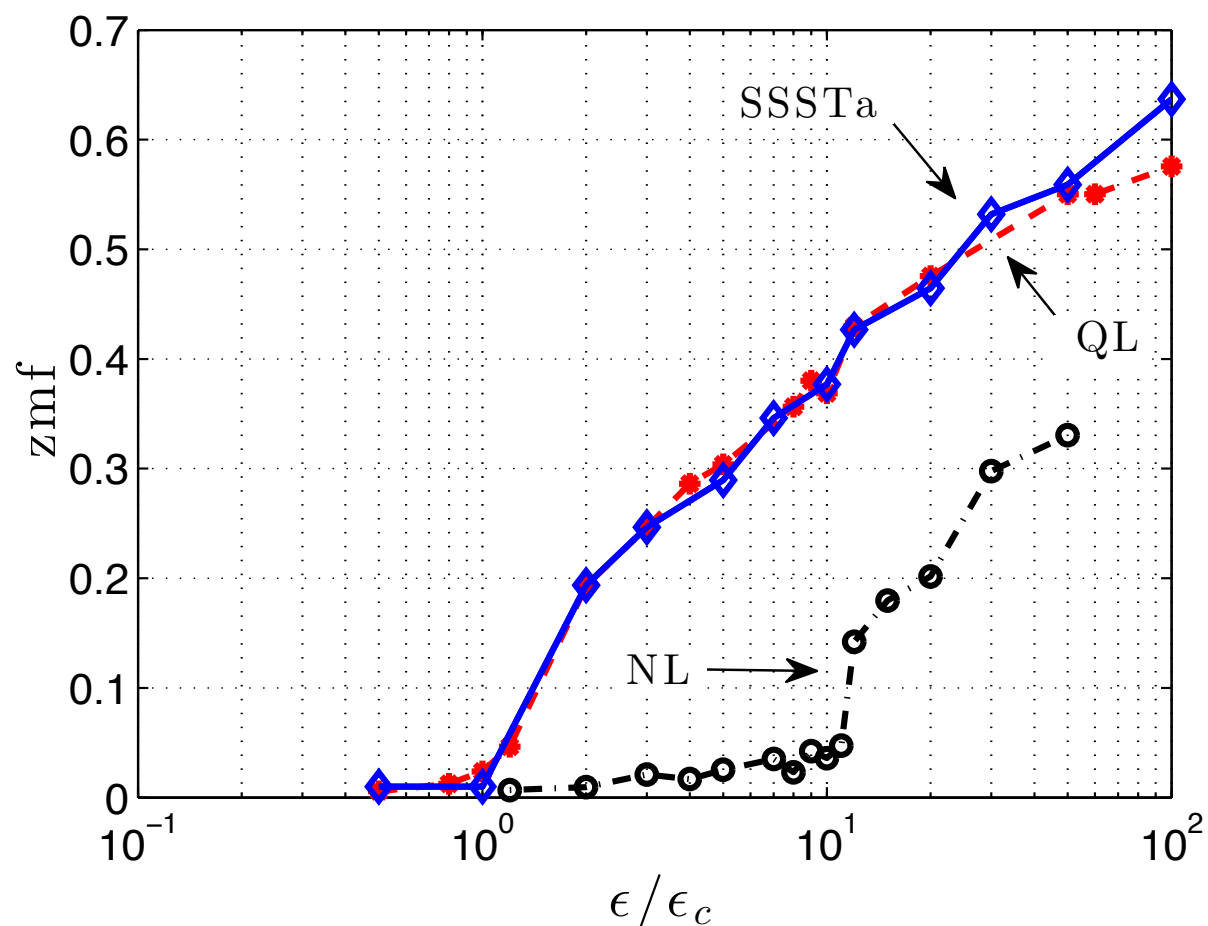
NIF, $r_m = r/10$



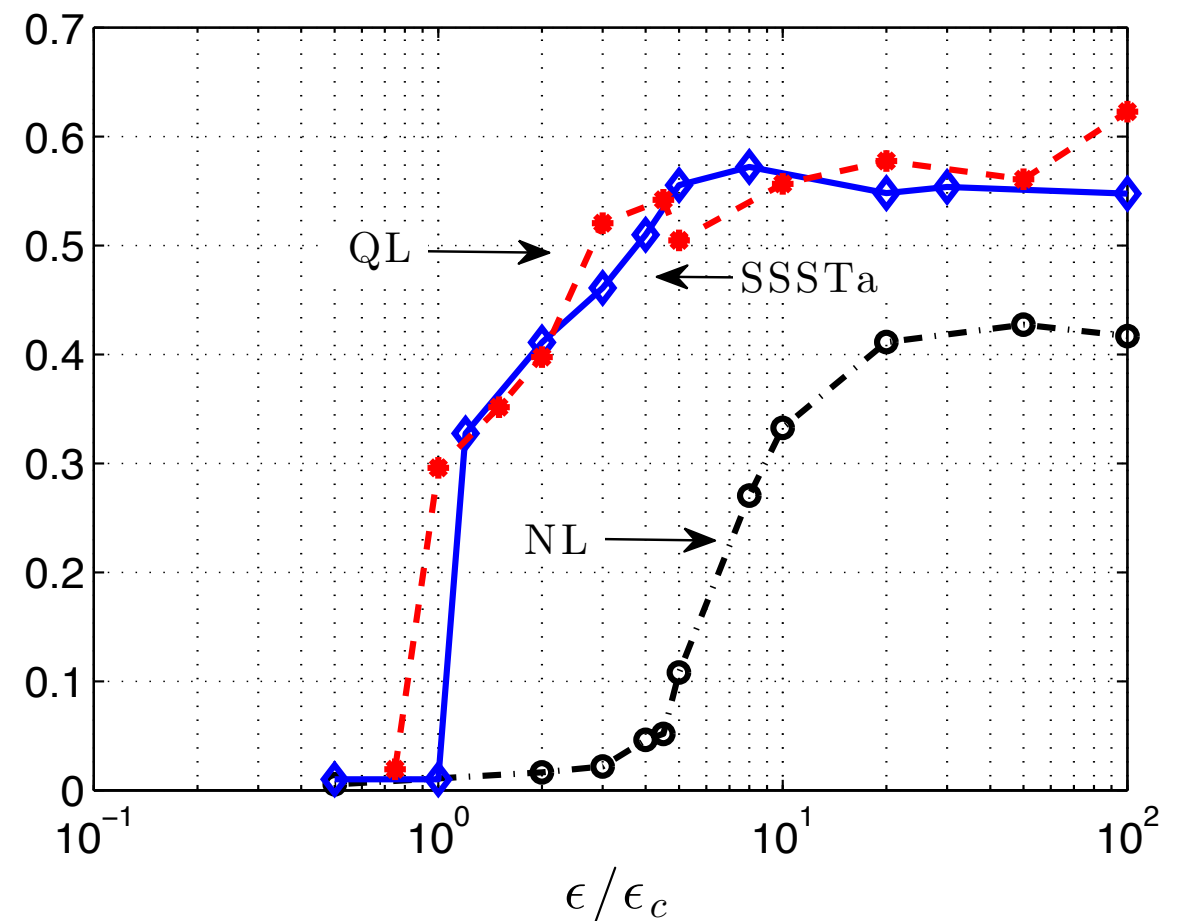


Bifurcation diagrams for $r = r_m$ (revisited)

NIF, $r = r_m = 10^{-2}$



IRFh, $r = r_m = 10^{-2}$

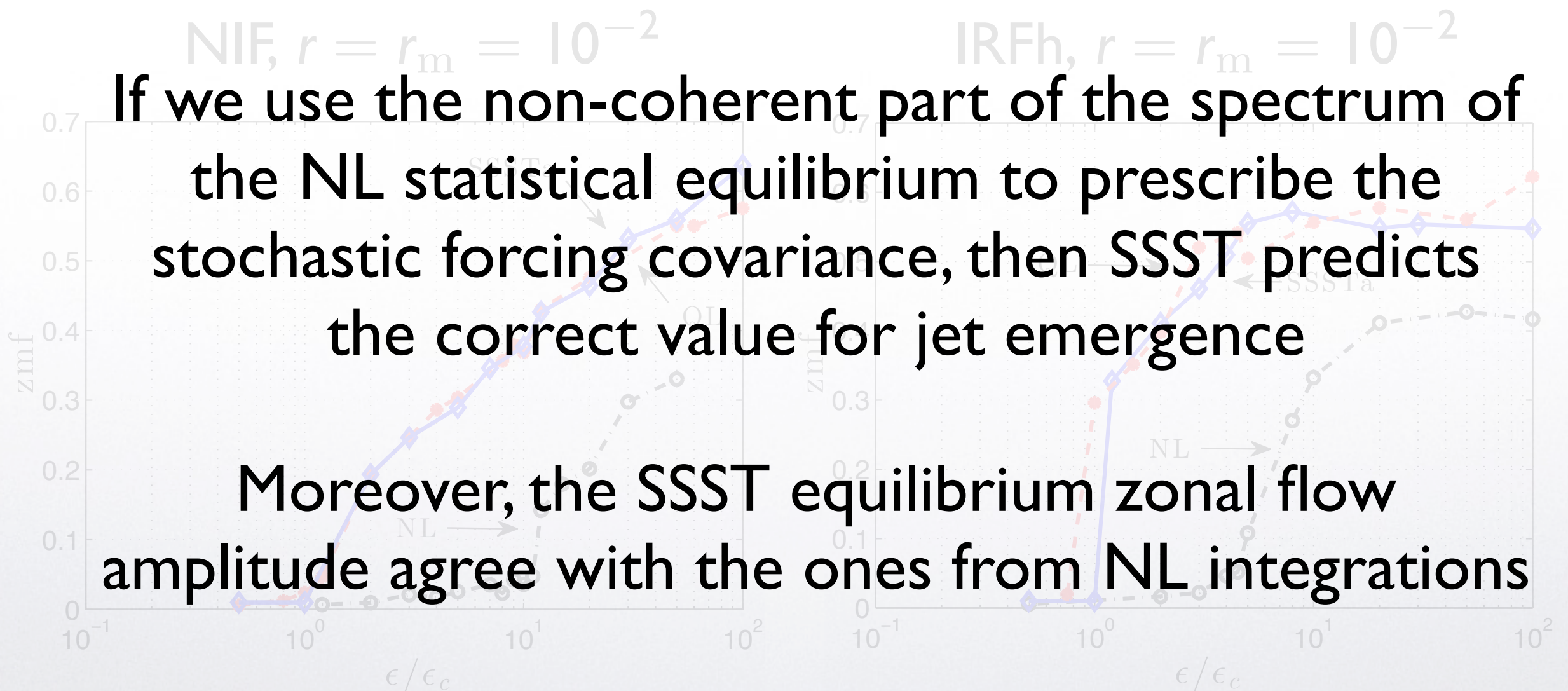




Bifurcation diagrams for $r = r_m$ (revisited)

If we use the non-coherent part of the spectrum of the NL statistical equilibrium to prescribe the stochastic forcing covariance, then SSST predicts the correct value for jet emergence

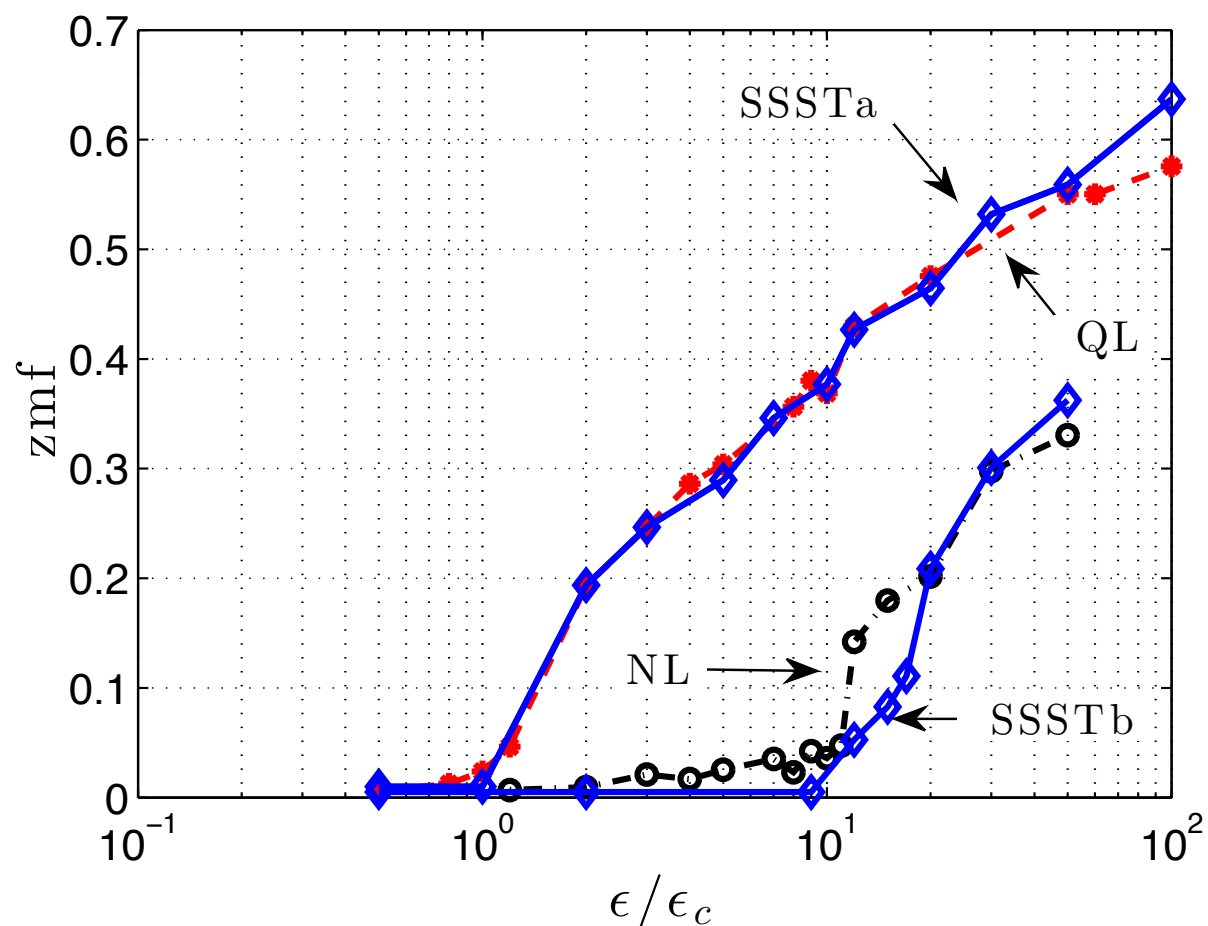
Moreover, the SSST equilibrium zonal flow amplitude agree with the ones from NL integrations



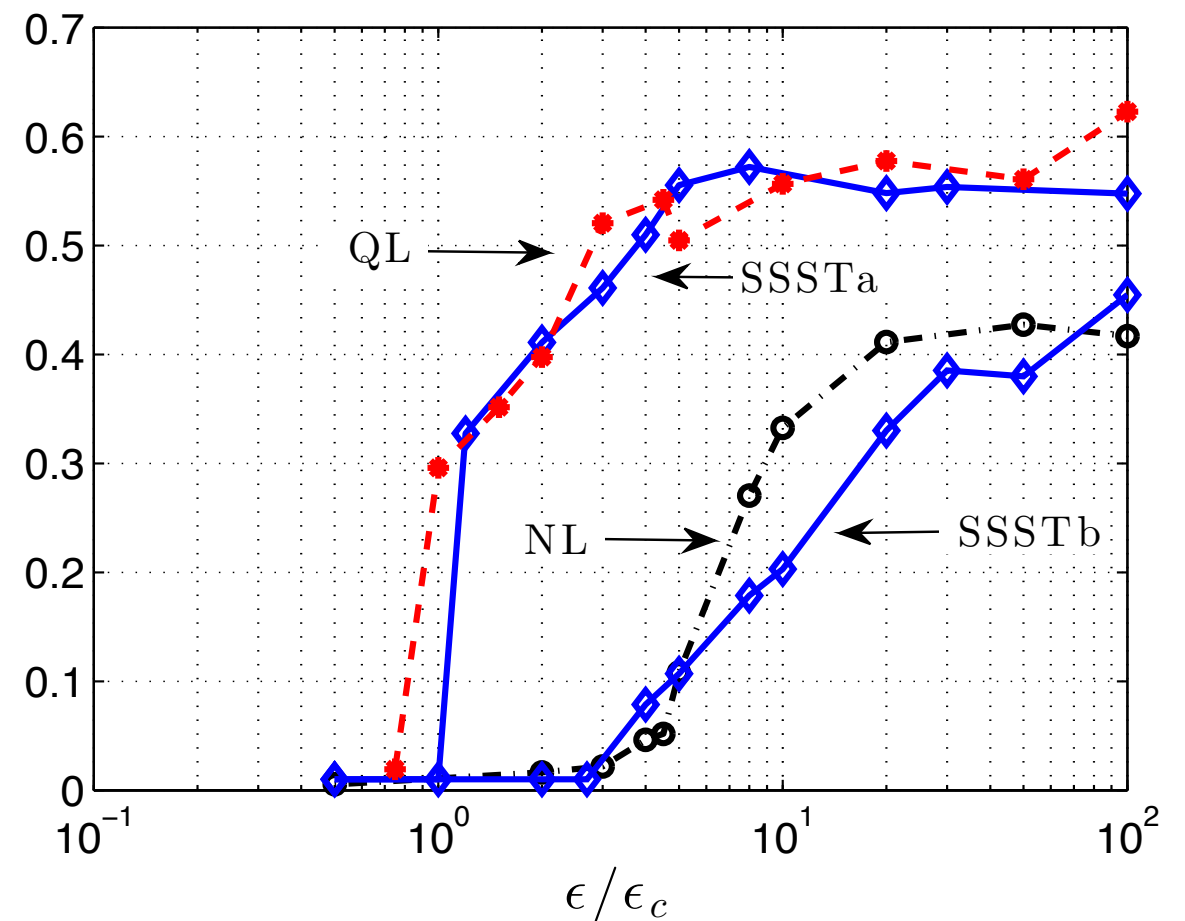


Bifurcation diagrams for $r = r_m$ (revisited)

NIF, $r = r_m = 10^{-2}$



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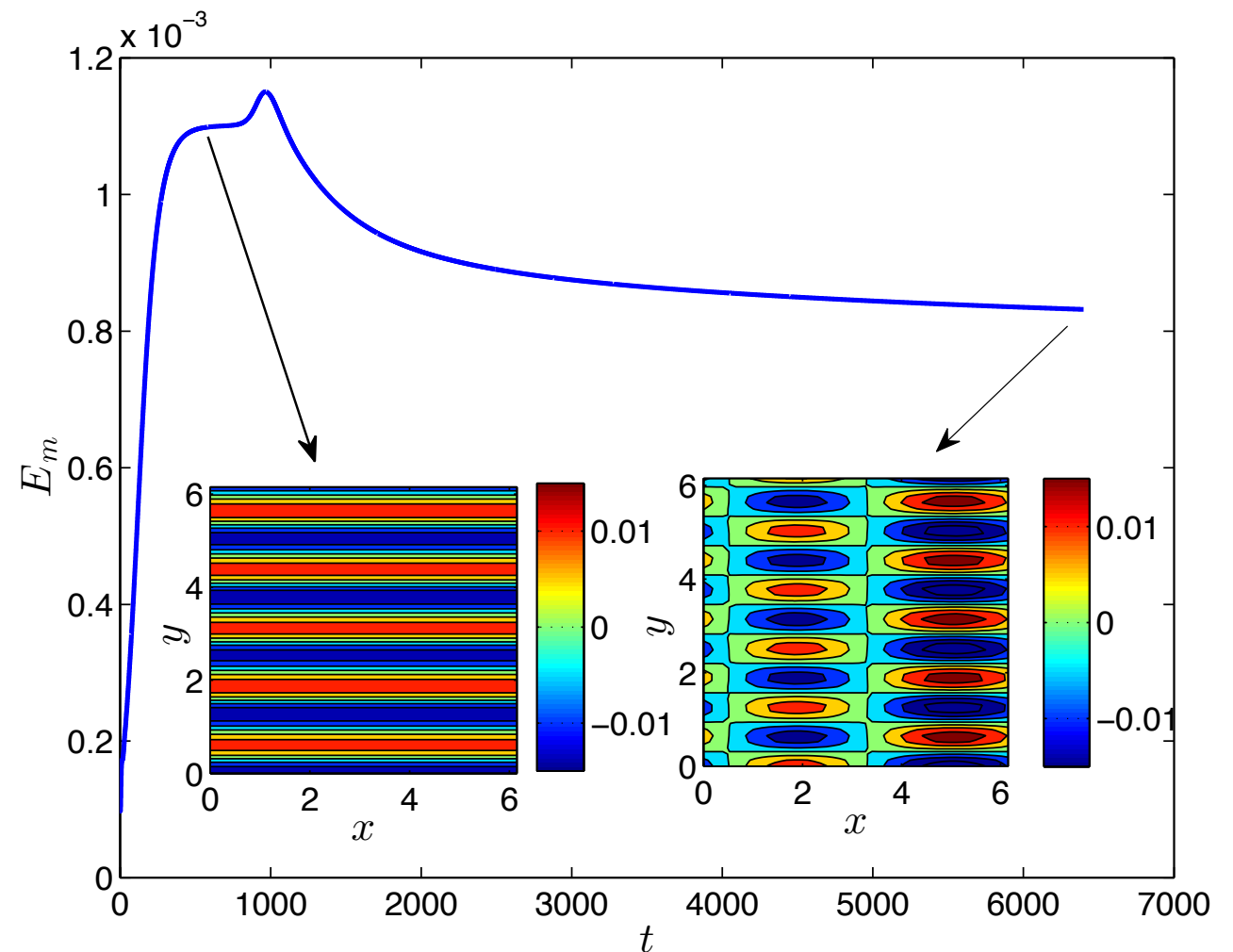


Generalized SSST admits non-zonal equilibria

IRFh at $\epsilon = 2\epsilon_c$

SSST at $\epsilon = 2\epsilon_c$ forms jets

generalized SSST initially goes to a jet equilibrium, which is SSST unstable and switches to a NZCS



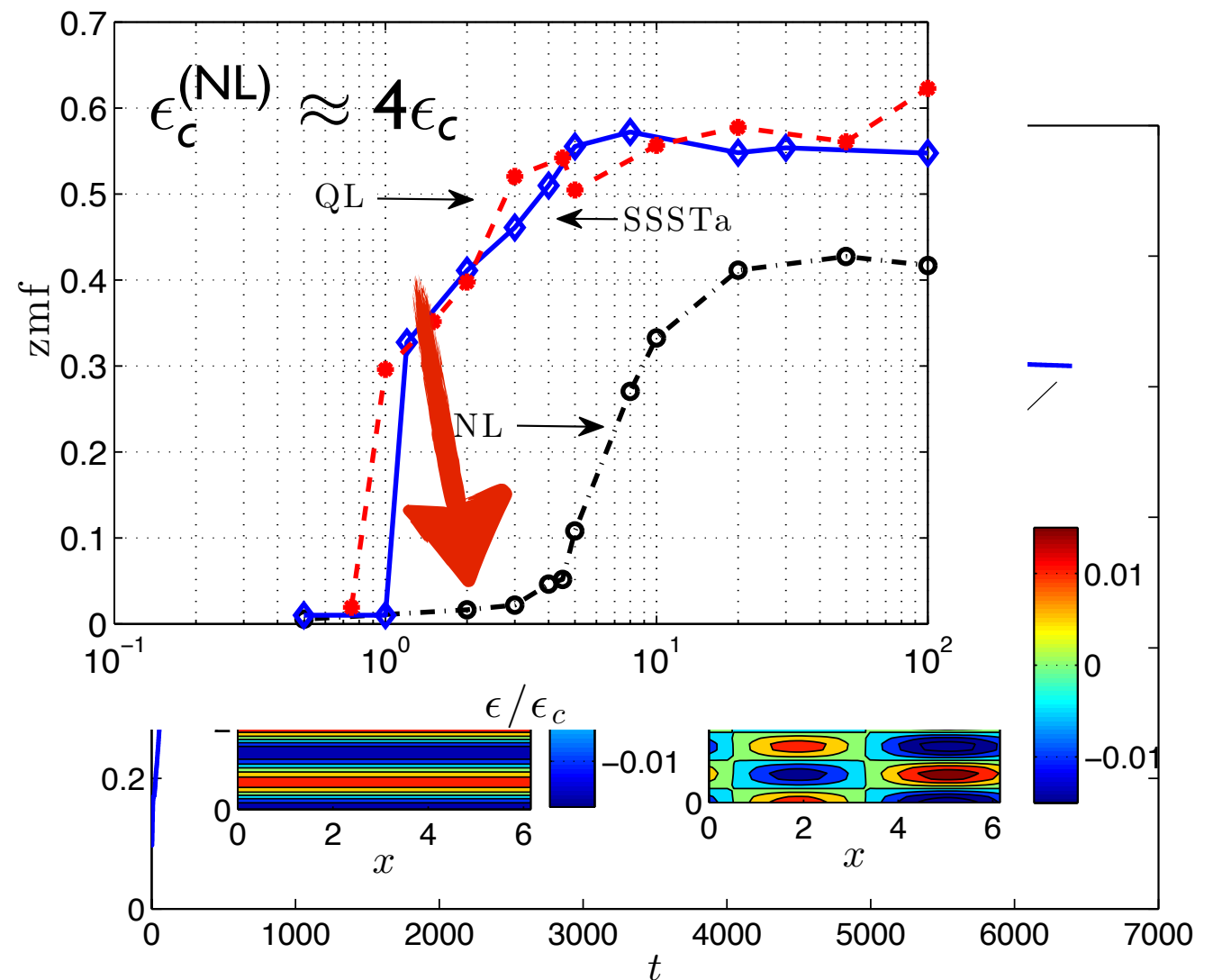


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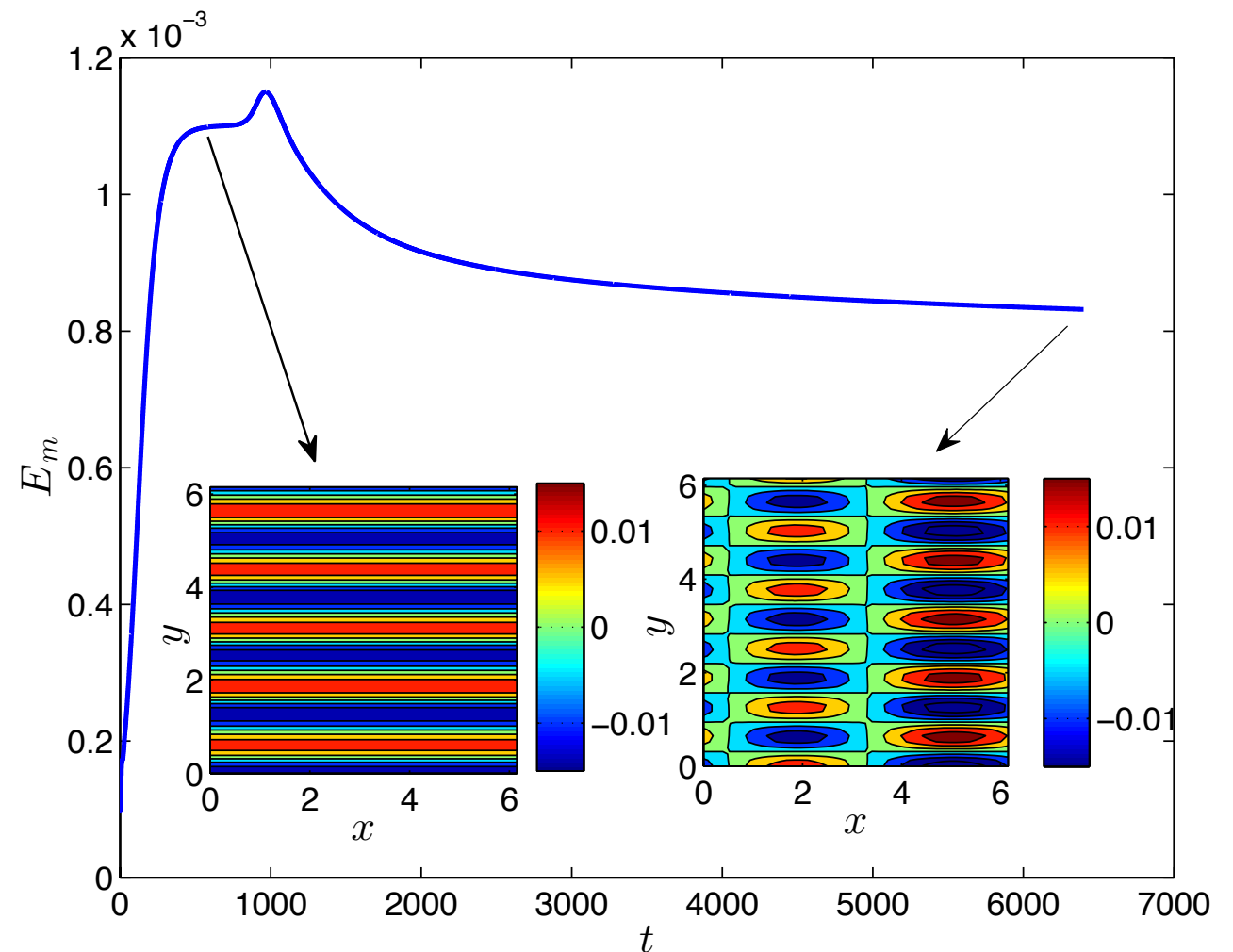


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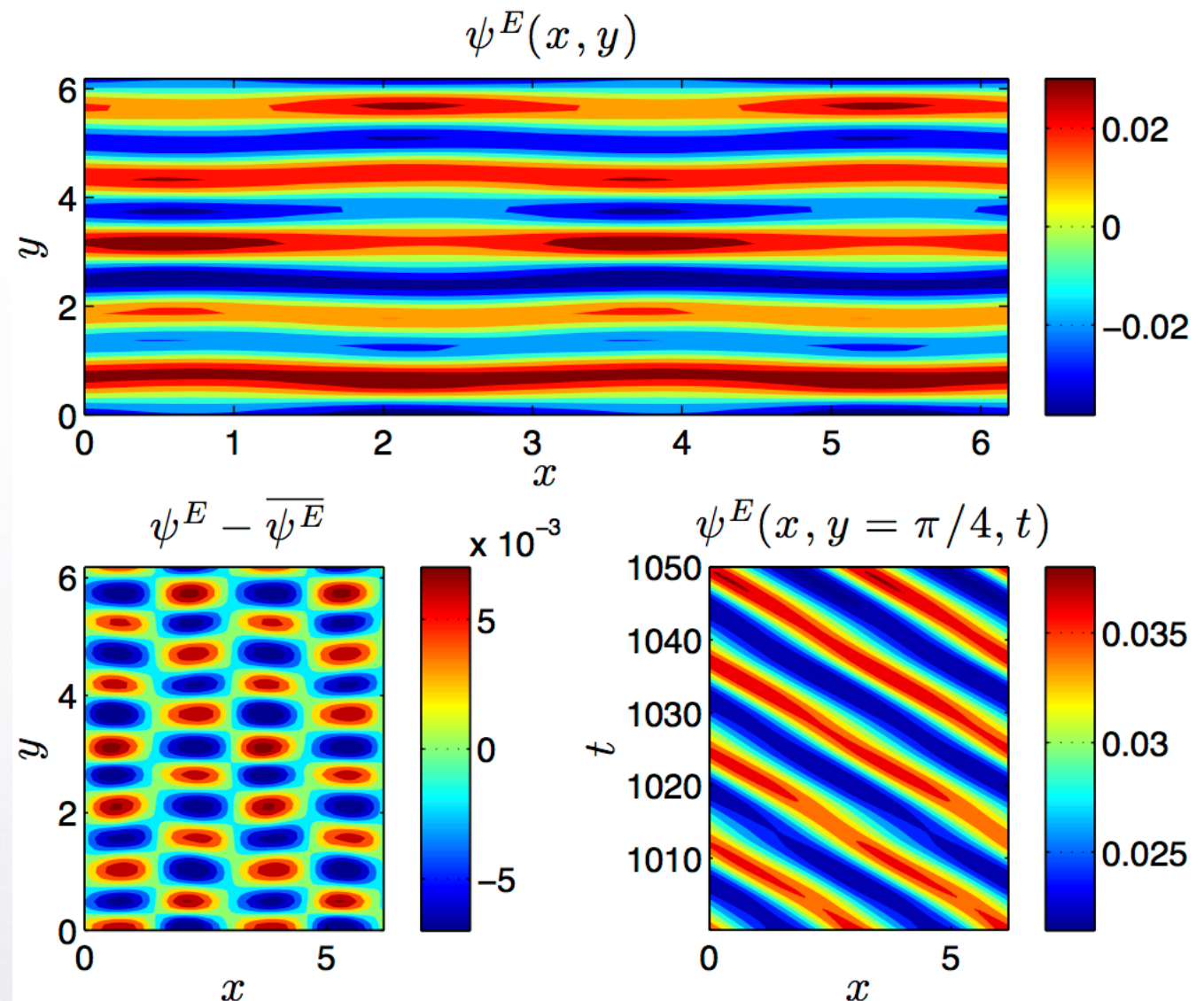




Generalized SSST equilibria show both zonal and non-zonal spectral peaks

IRFh at $\epsilon = 6\epsilon_c$

generalized SSST admits equilibria with zonal as well as non-zonal spectral components

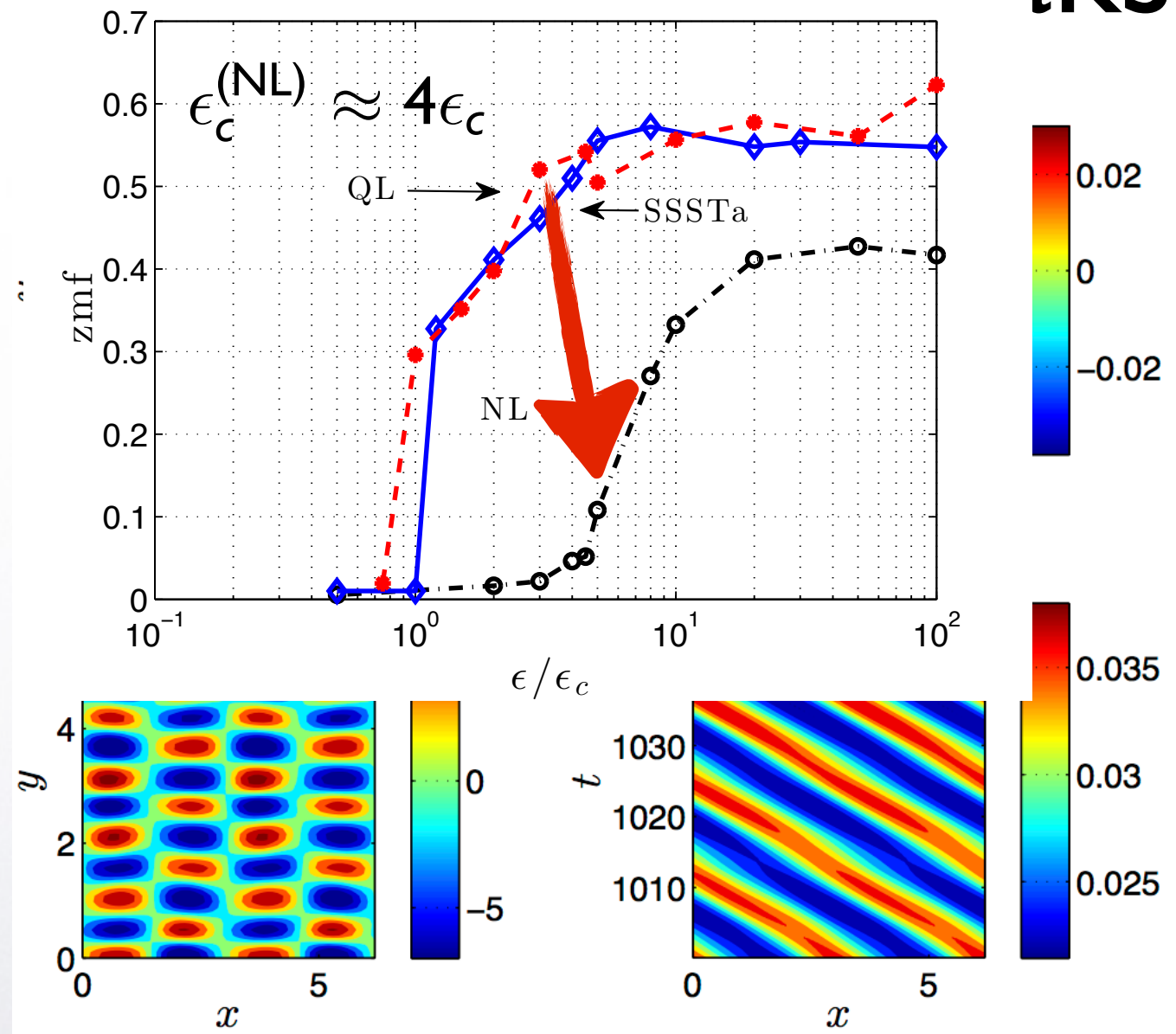




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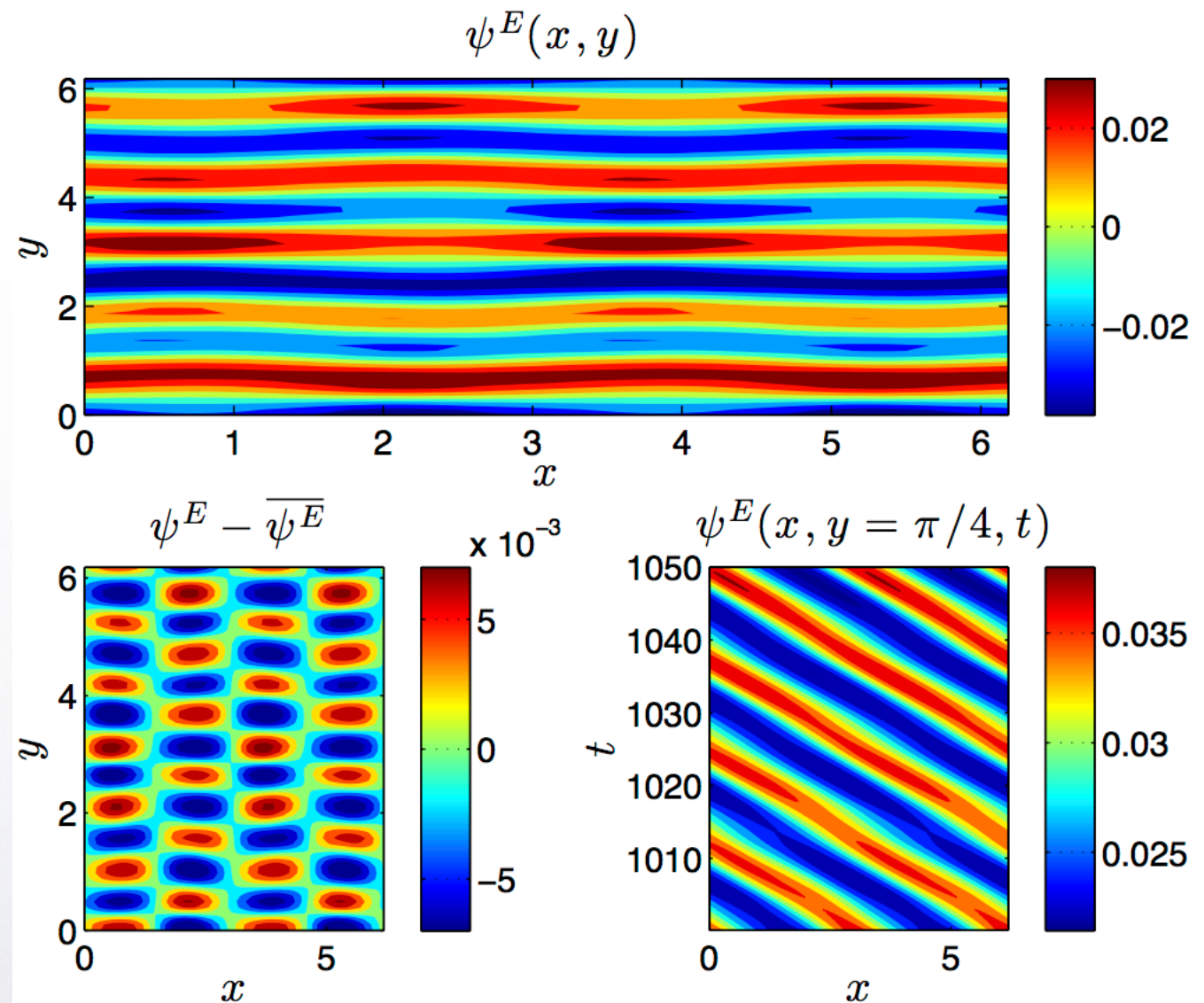




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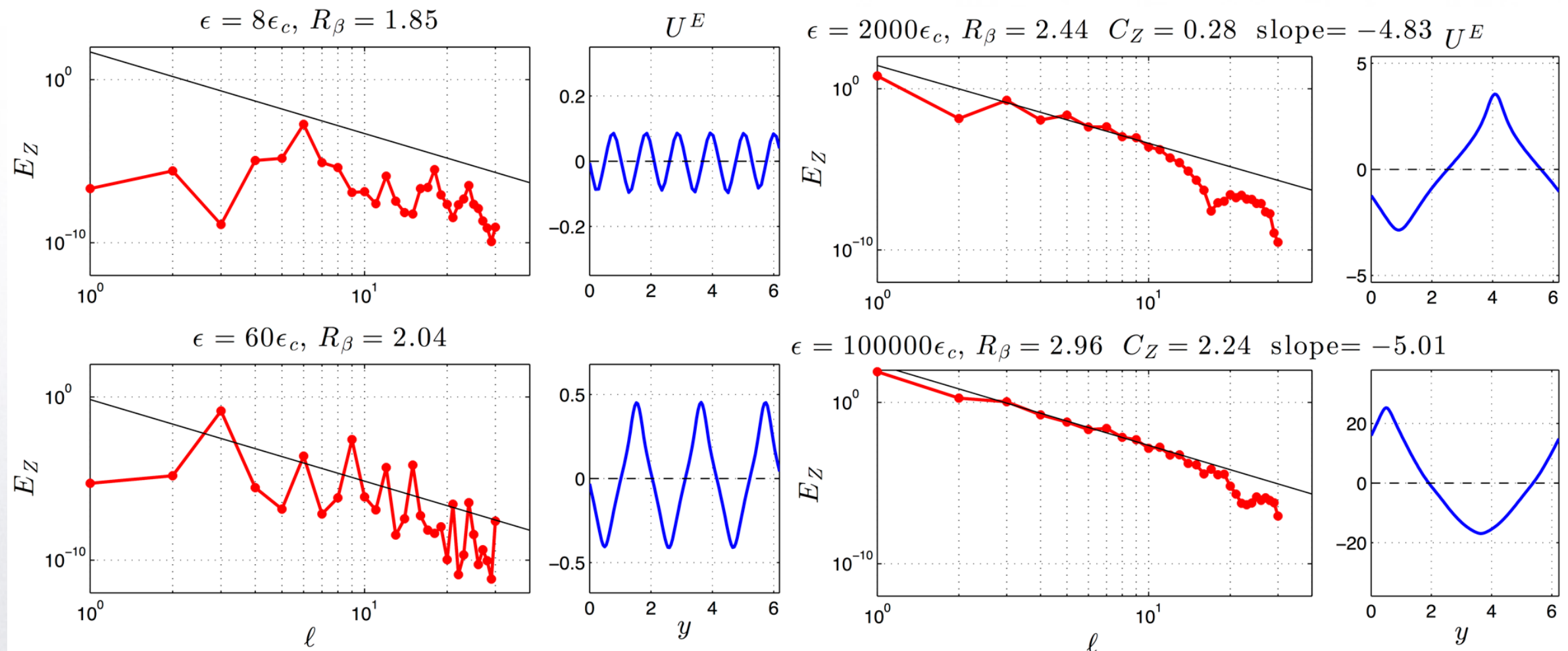
Things to discuss

- ▶ What is the shape of the SSST mean flow equilibria? Does it capture the NL jet shape?
- ▶ Jet mergings / why do they occur?
(is it because of U becomes hydrodynamically unstable?)
- ▶ PV staircases (?)



SSST equilibrium zonal flow spectral structure

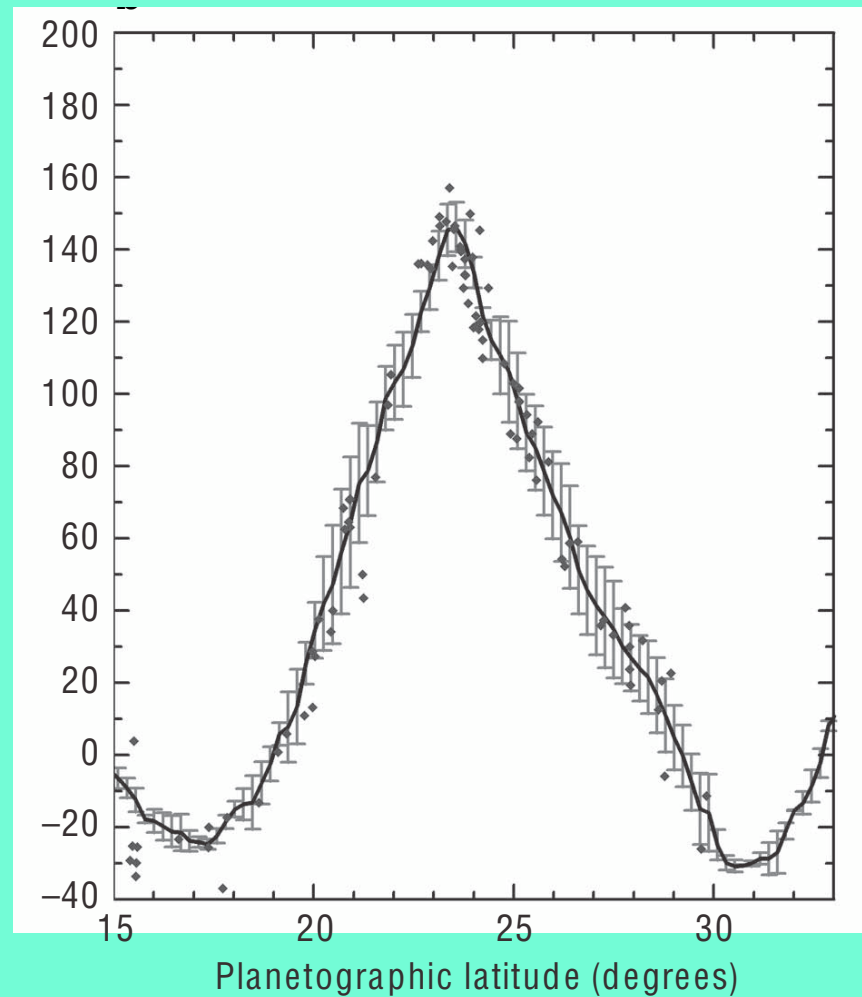
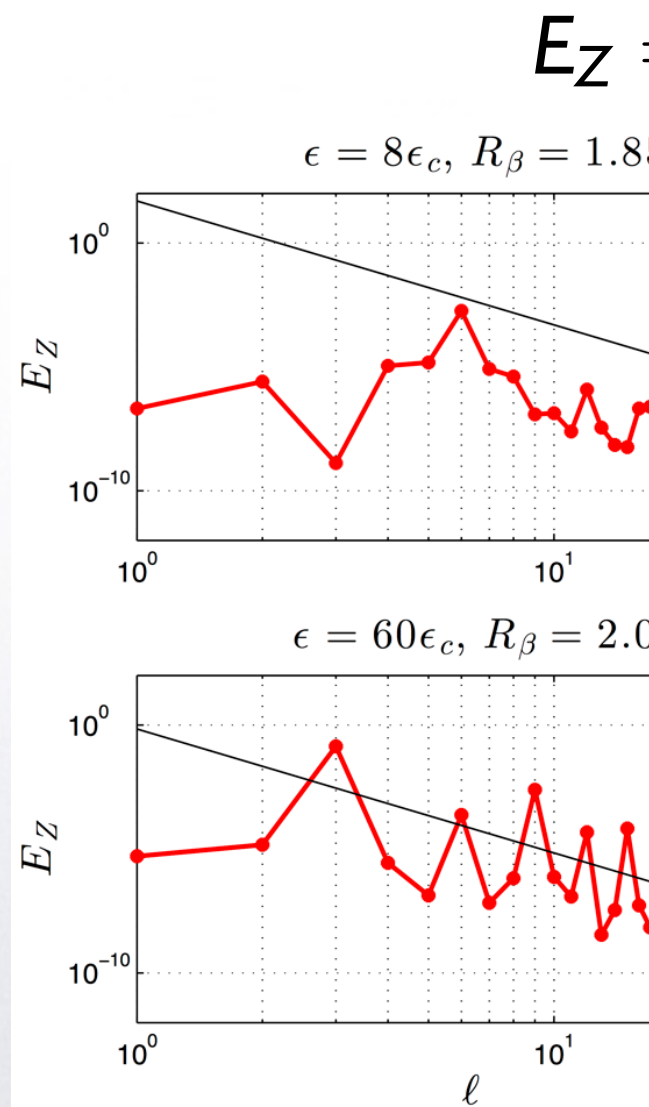
$$E_Z = E(k=0, \ell) = C_Z \beta^2 \ell^{-5}, C_Z \approx 0.5$$



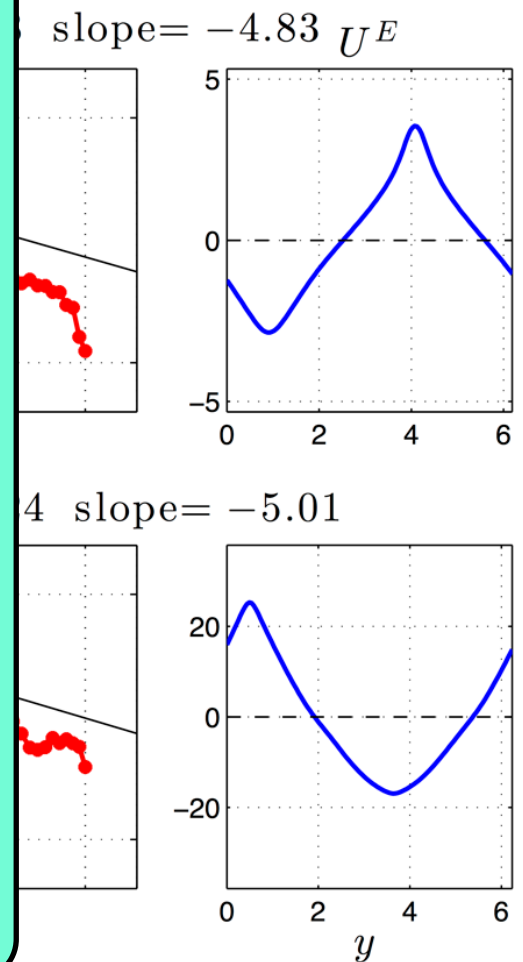


SSST equilibrium zonal flow spectral structure

Jupiter's 23N jet



(Sanchez-Lavega et. al., 2008)



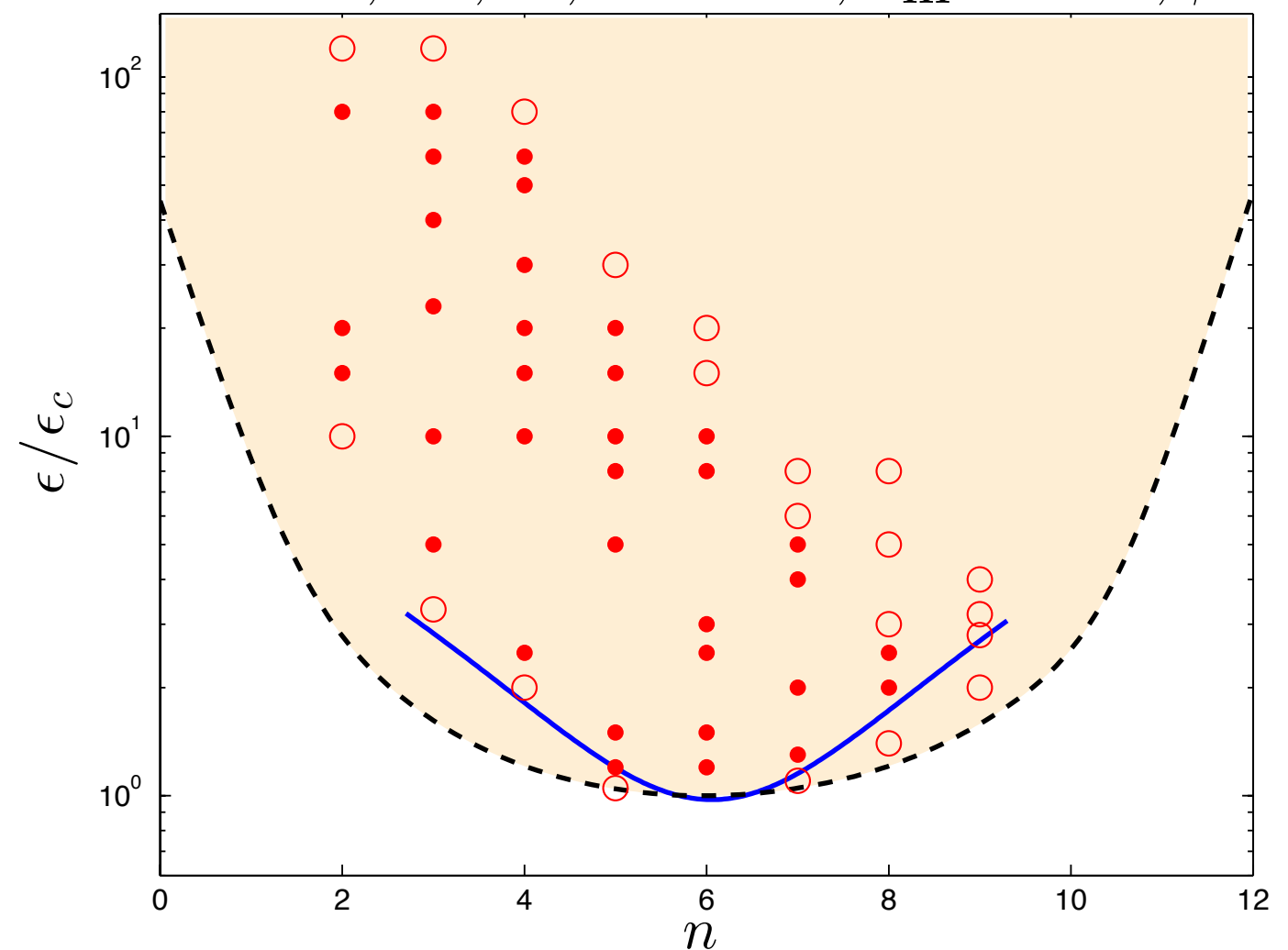


Jet merging is due to SSST instability

SSST equilibria
(stable & unstable)
are hydrodynamically stable

SSST instability comes prior
to hydrodynamic instability
as we move upwards in a
constant n axis

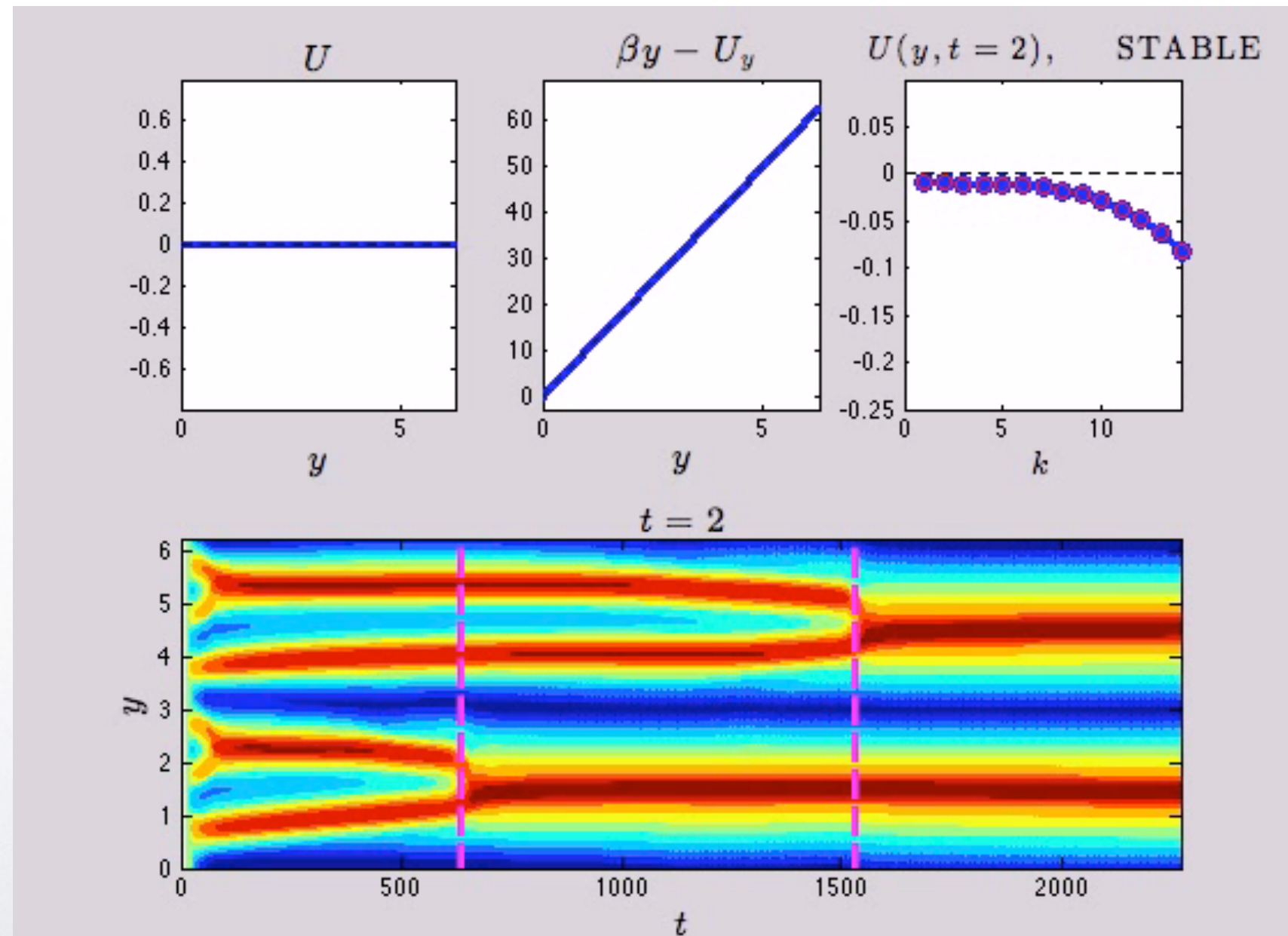
NIF at $k = 1, \dots, 14$, $r = 10^{-1}$, $r_m = 10^{-2}$, $\beta = 10$





Jet merging is due to SSST instability

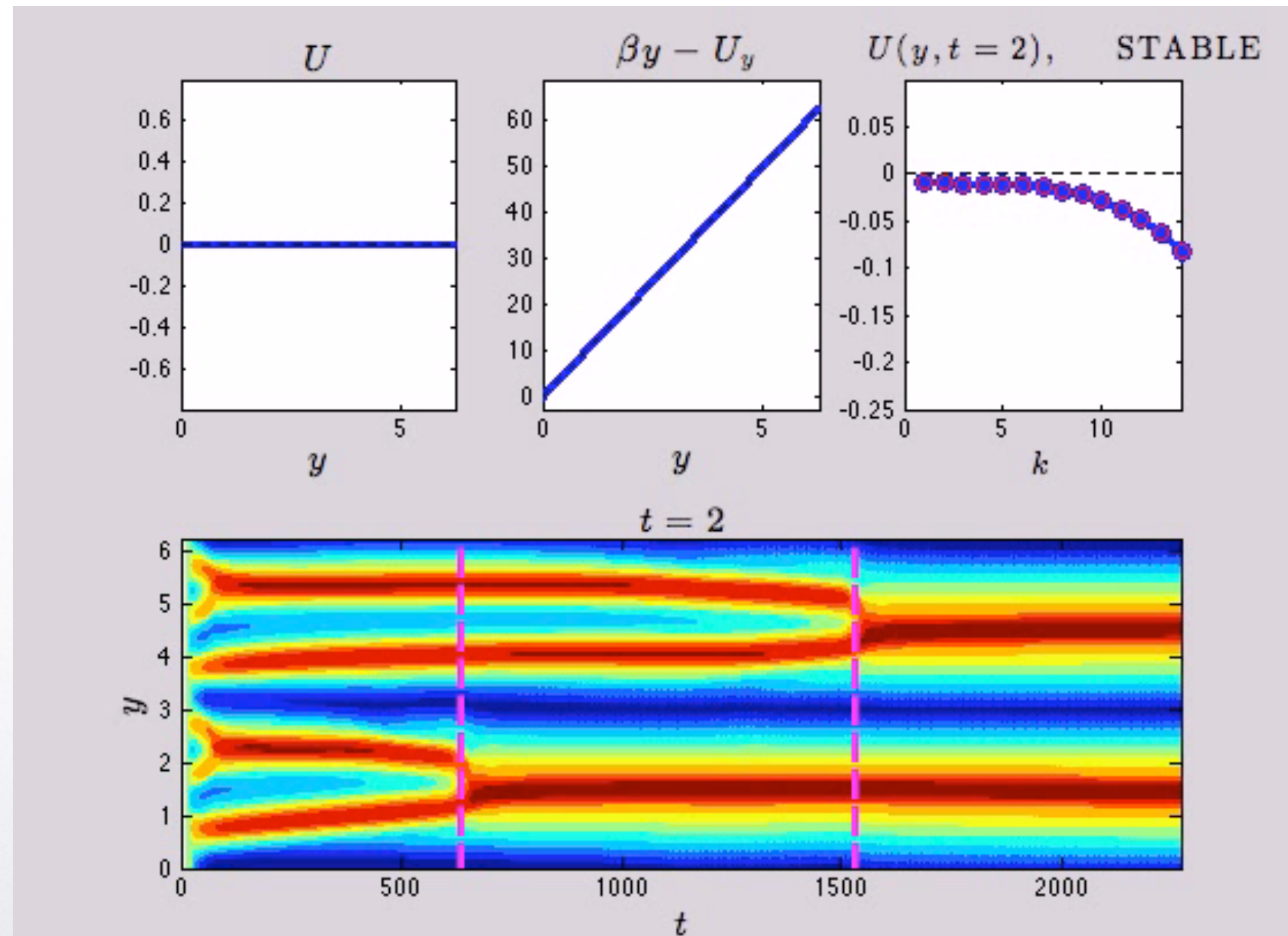
NIF at $k = 1, \dots, 14$
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 $\epsilon = 100\epsilon_c$





Jet merging is due to SSST instability

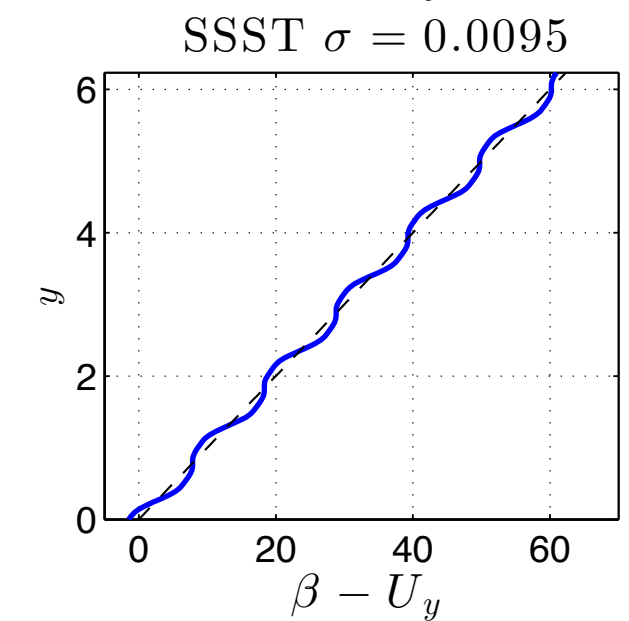
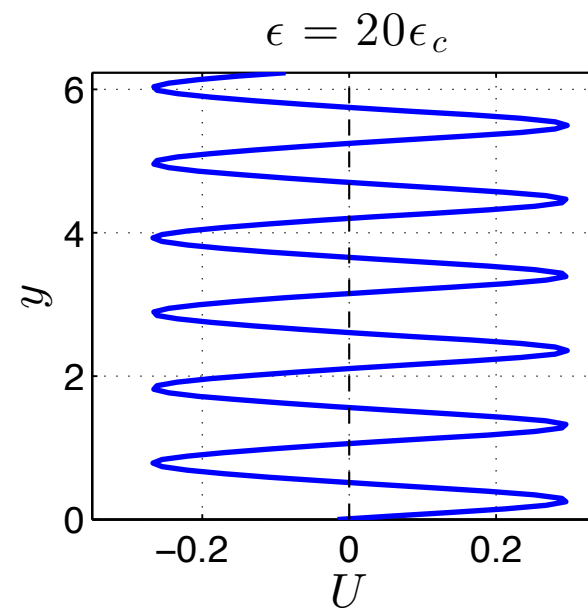
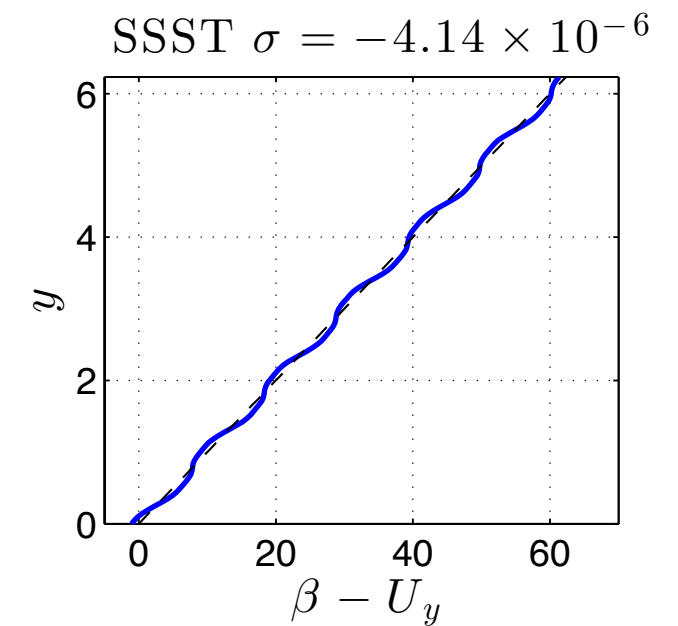
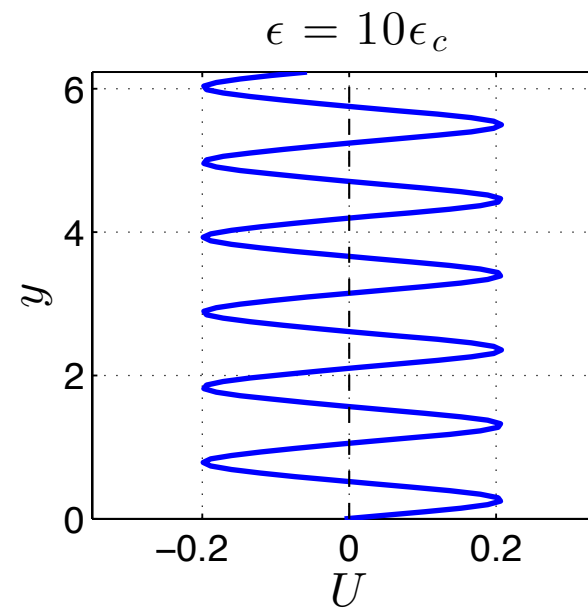
NIF at $k = 1, \dots, 14$
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 $\epsilon = 100\epsilon_c$





PV Staircases

NIF, $k = 1, \dots, 14$
 $r = 0.1$, $r_m = r/10$, $\beta = 10$

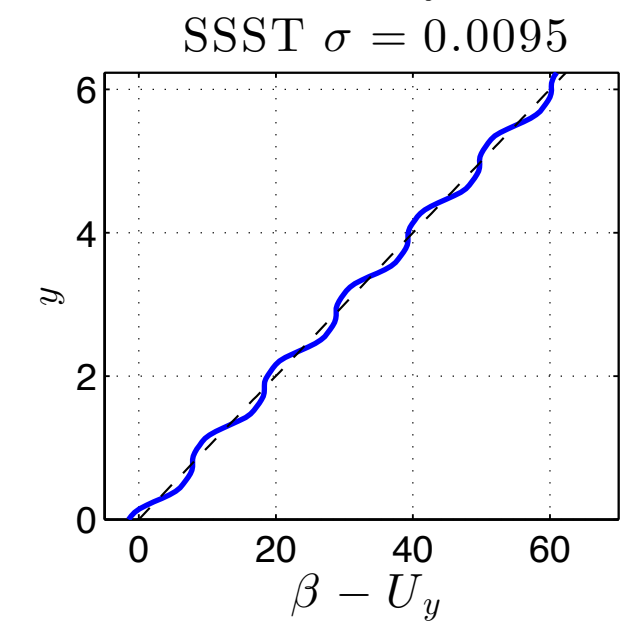
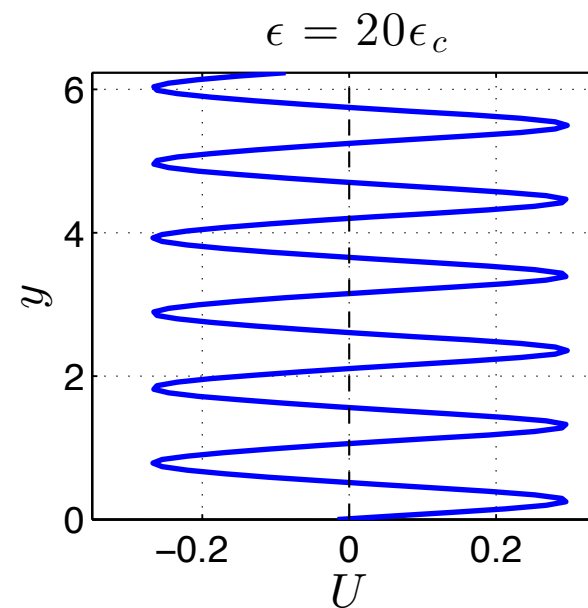
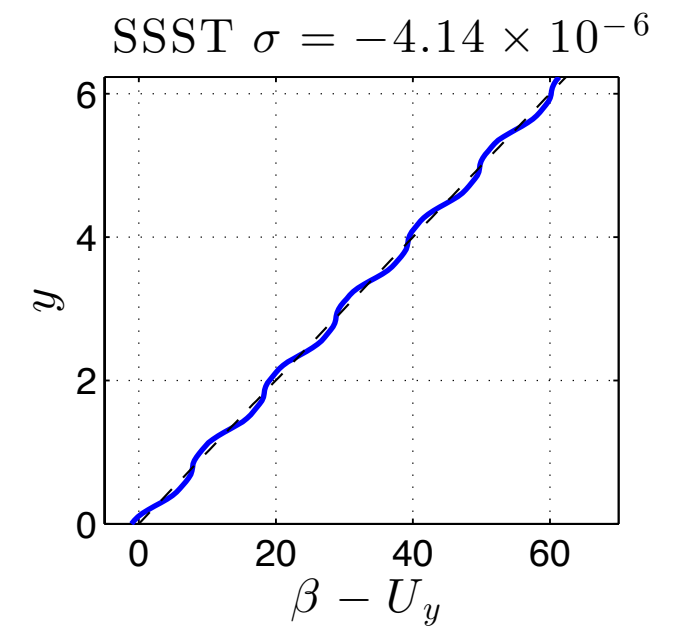
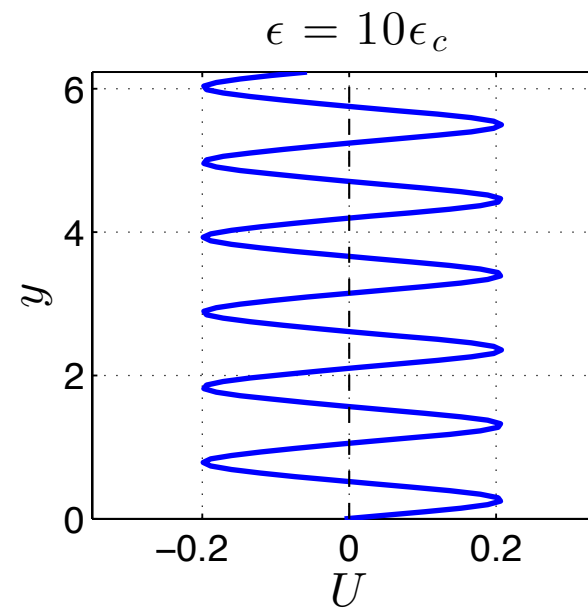
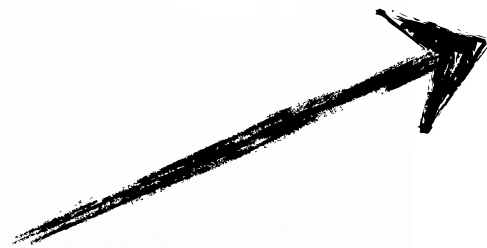




PV Staircases

NIF, $k = 1, \dots, 14$
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SSST stable

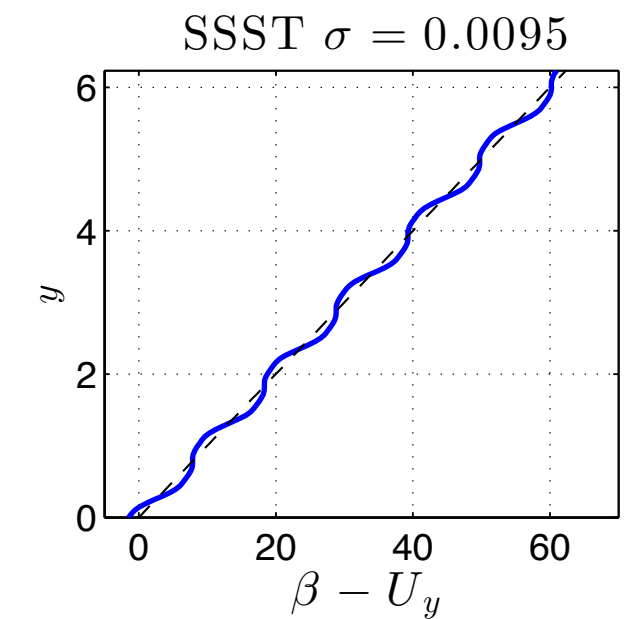
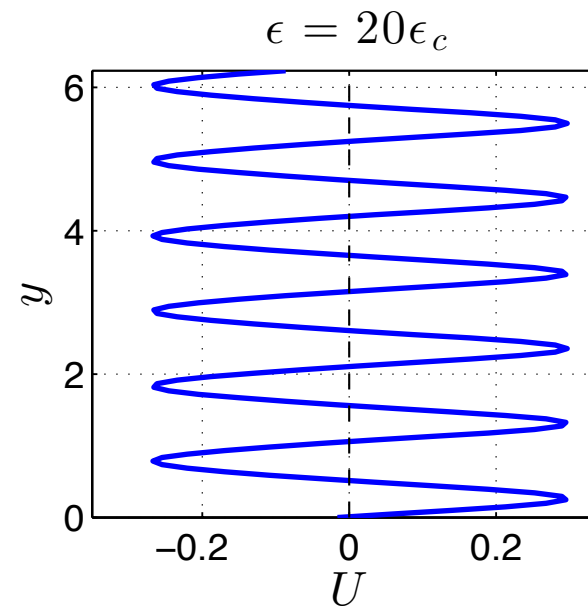
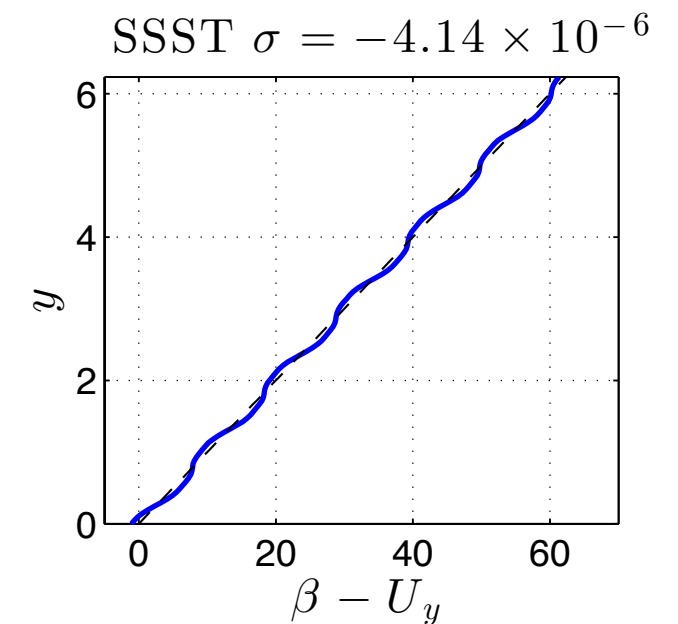
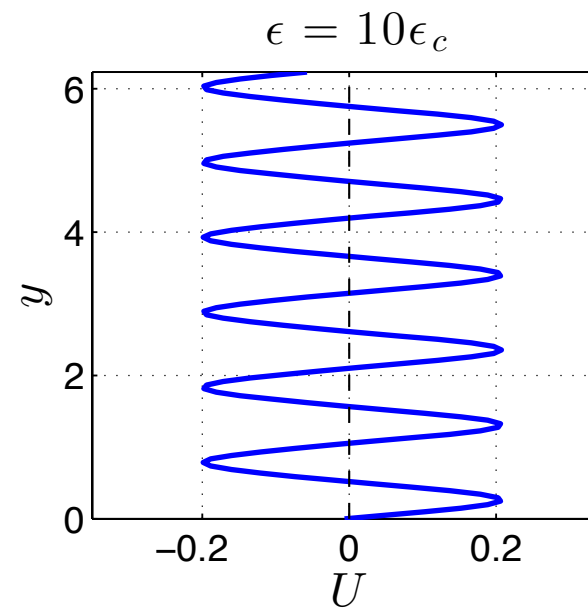




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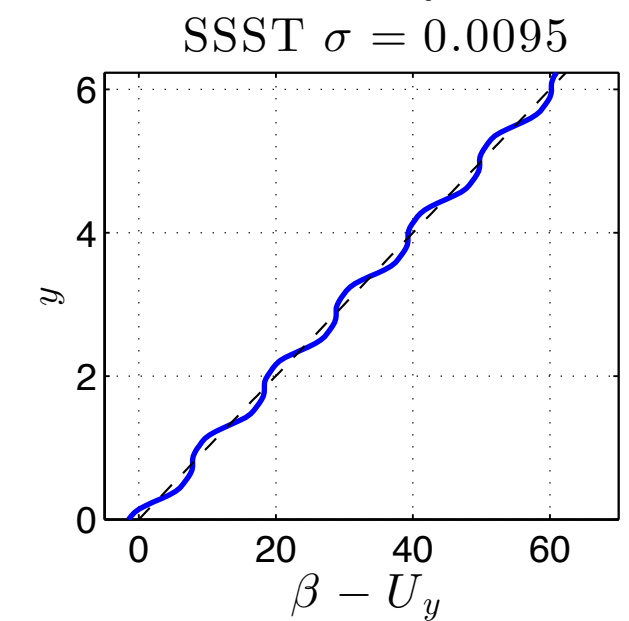
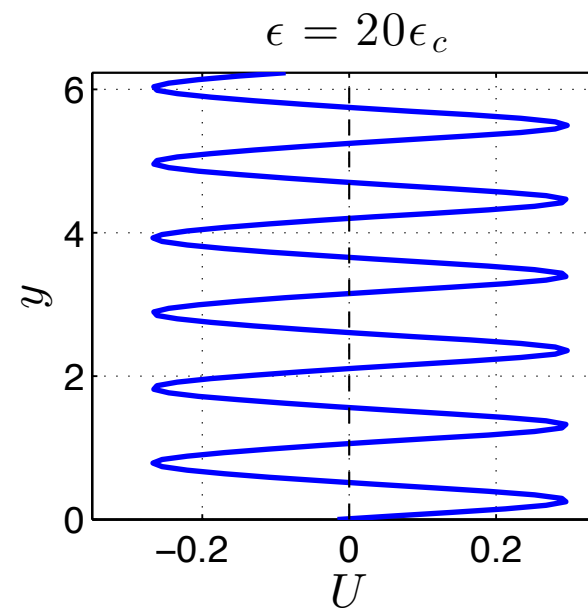
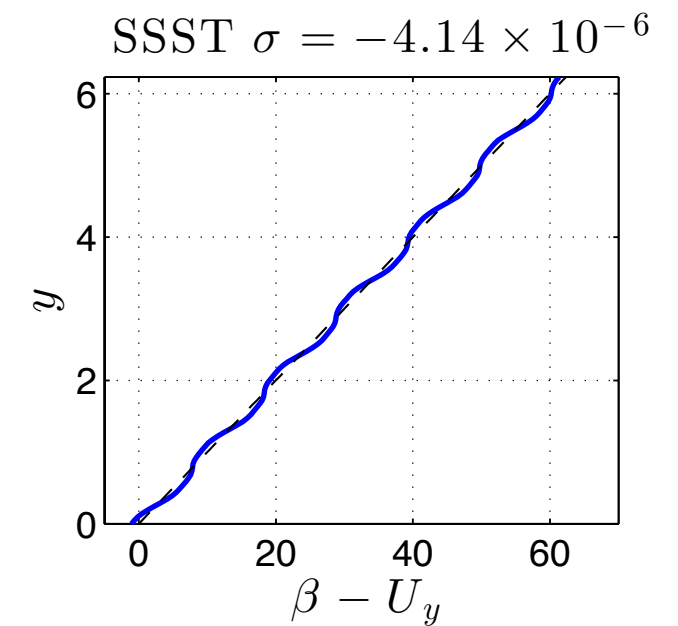
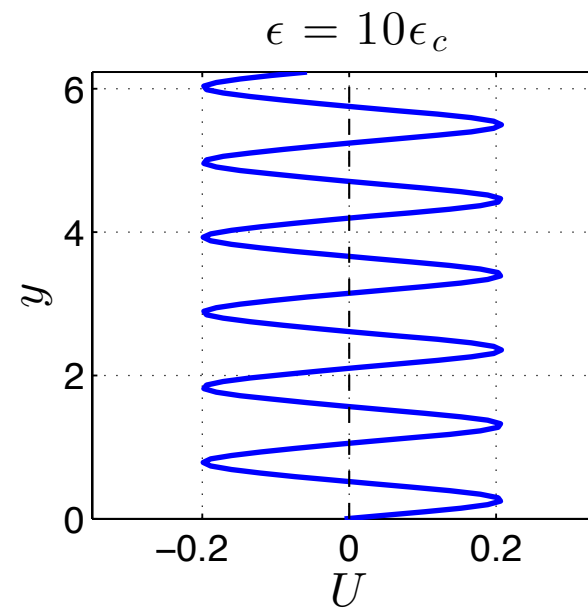
SSST unstable





PV Staircases

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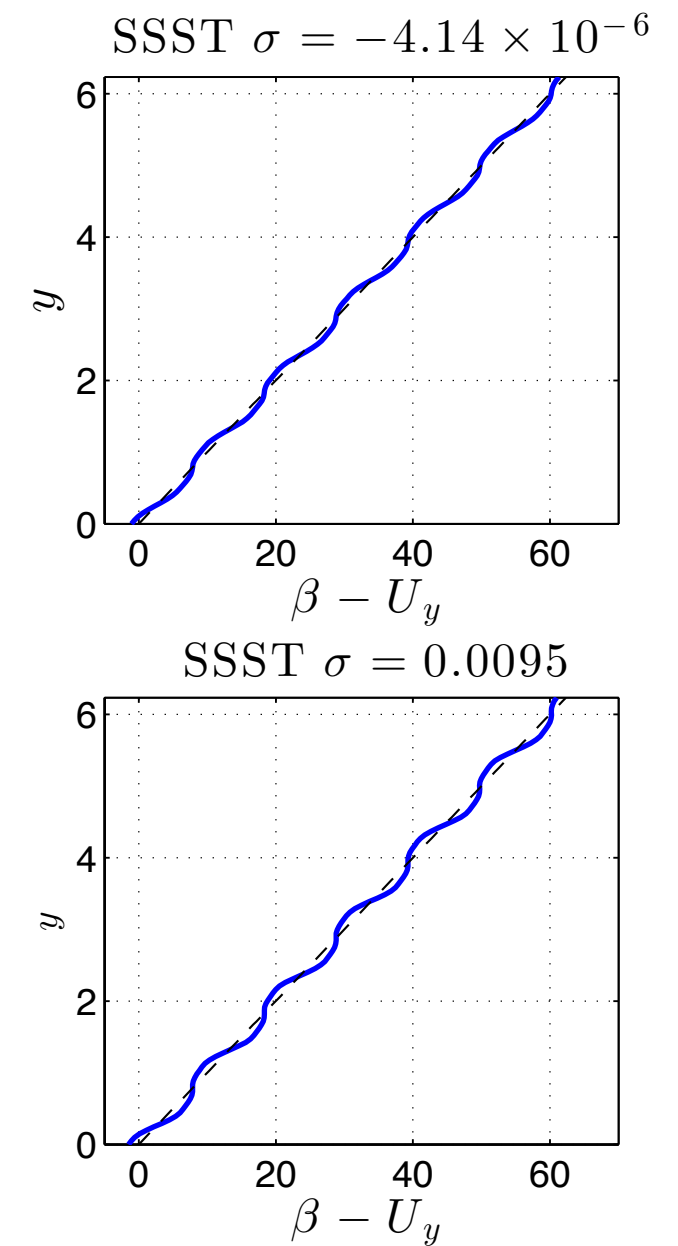
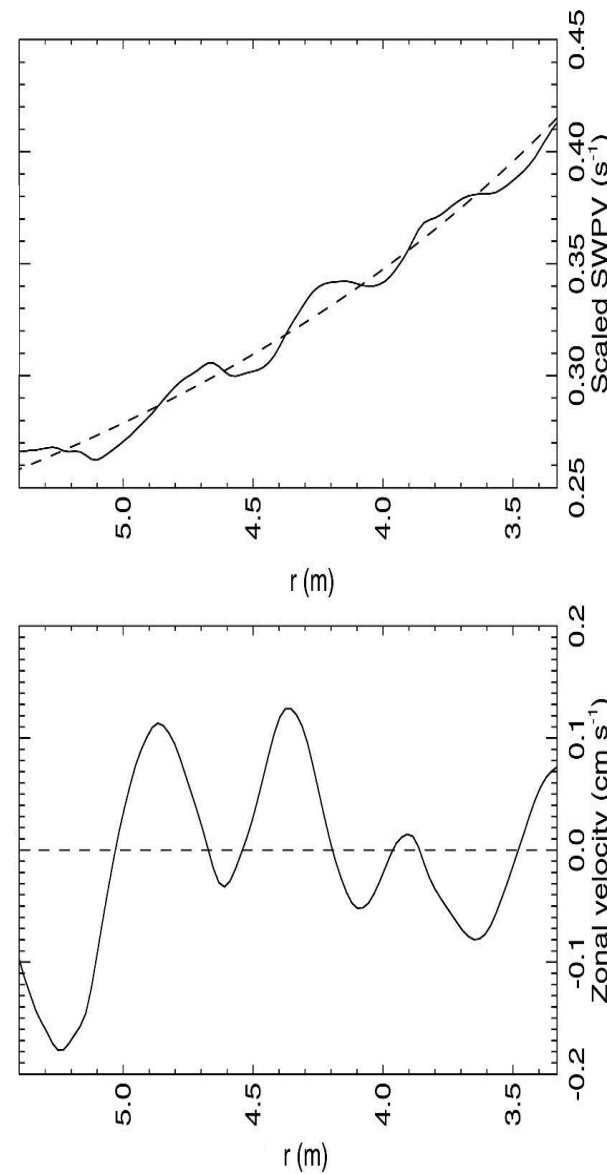


PV Staircases

NIF, $k = 1$,
 $r = 0.1$, $r_m = r/$

PV staircases in
rotating tank
experiment

(Read et. al., 2007)





Conclusions



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- with the modification of the turbulent spectrum due to the emergence of NZCS, SSST captures the bifurcation for jet emergence
- stochastic excitation of SSST eigenmodes reveals that these modes underlie in the dynamics even in subcritical cases



Thank you

This work has been
supported by



Constantinou, N.C, Ioannou, P.J. and Farrell, B.F., 2012:
Emergence and equilibration of jets in beta-plane turbulence.
(arXiv:1208.5665 [physics.flu-dyn])